

2.3 Symmetry and Classical Field Theory

- The ACTION $S = \int d^d x \mathcal{L} = \int d^d x \times \mathcal{L}$ describes the whole of classical dynamics
- Any solution to classical e.o.m. is a turning point in S e.g. local maximum
- Want S to be invariant under physical symmetry
 $\Leftrightarrow$ symmetries map solⁿ. to e.o.m. to other related (same energy) solⁿ.

2Classical Field Theory
SYMMETRY (cont)

Symmetry only mixes objects (field, particles, ...) which lie in same irreducible representation of the group.
(and vice-versa irrep $\Rightarrow$ group of related objects)

Lorentz/Poincaré group labels irreps by

a) MASS, m^2 or $\mathbb{R}^+$

b) SPIN, S , $0, \frac{1}{2}, 1, \dots$

e.g. Two degrees of freedom of photon (polarisation) linked in one "particle" because space-time symmetry mixes them ($m=0, S=1$)
Three degrees of freedom of π^0 (Long, + 2x Transverse) (since $m > 0, S=1$)

) Relativistic Space-Time Symmetry - Poincaré Group

(i) Translations $x^\mu \rightarrow x'^\mu = x^\mu + a^\mu$
 $\uparrow$ constant

(ii) Rotations/Boosts $x^\mu \rightarrow x'^\mu = \Lambda^\mu_\nu x^\nu$

No detailed study here, only results quoted

Ex. Relativistic Scalar Field Theory $x^\mu = \Lambda^\mu_\nu x'^\nu$ Lorentz subgroup

$$S = \int dt L = \int d^4x \mathcal{L}$$

EFS • $\int d^4x = \int d^4x'$ invariant (EFS)

$\Rightarrow$ Work with Lagrangian density

• K.E. $\dot{\phi}^2 = (\partial_t \phi)^2$ NOT invariant

But we have $\left(\frac{\partial \phi}{\partial t}\right)^2 - c^2 \left(\frac{\partial \phi}{\partial x}\right)^2$ in 1D example

c=1?
EFS

$$\Rightarrow (\partial_\mu \phi)(\partial^\mu \phi) = \left(\frac{\partial \phi}{\partial t}\right)^2 - \left(\frac{\partial \phi}{\partial x}\right)^2 \dots \text{(with c=1 here!)} \quad \text{Active!}$$

This is invariant IFF $\phi(x) \xrightarrow{1} \phi'(x') = \phi(\Lambda^{-1} x)$
i.e. coordinates change but function is UNCHANGED/INVARIANT

This is CALLED a (LORENTE) SCALAR FIELD
c.p. $A^\mu(x) \rightarrow \Lambda^\mu_\nu A^\nu(x) = \Lambda^\mu_\nu A^\nu(\Lambda^{-1} x)$

Equivalent to having ZERO SPIN

Tony 1.27) • Potential Terms (= No derivatives)

PI2 $\int d^4x \phi^2(x) \xrightarrow{\text{Transform}} \int d^4x' (\phi'(x'))^2$ (ϕ unchanged)

$$= \int d^4x' (\phi(x'))^2$$

$$= \int d^4x (\phi(x))^2 \quad \downarrow \text{change label}$$

$\Rightarrow$ Action invariant

⇒ our 1D toy model has given us
a generic form of a relativistic scalar (=L.inv)
(spin 0) field theory

Action

Lagrangian Density

$$S = \int d^4x \mathcal{L}, \quad \mathcal{L} = \frac{1}{2} \left\{ (\partial_\mu \phi)(\partial^\mu \phi) - m^2 \phi^2 \right\}$$

↑
convention

↑
assert: MASS² of particle
($c=\hbar=1$)
(now)

if $\phi(x) \in \mathbb{R} = \text{ONE degree of freedom}$

e.o.L4, 12/10/14

(See P.S. ch2, 2 p16-17)

For later use, we find the Hamiltonian density $\mathcal{H}$

Momentum Density $\Pi(x) := \frac{\partial \mathcal{L}}{\partial \dot{\phi}} = \dot{\phi}, \quad \mathcal{L} = \int d^3x \mathcal{L}$

⇒ $\mathcal{H} = \Pi \dot{\phi} - \mathcal{L}$ where $\mathcal{H} = \int d^3x \mathcal{H} = \text{Hamiltonian}$

$= \Pi^2 - \frac{1}{2} \left\{ \dot{\phi}^2 - (\nabla \phi)^2 - m^2 \phi^2 \right\}$

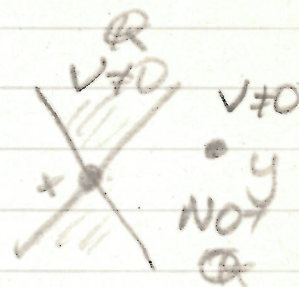
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p32)

⇒ $\mathcal{H} = \left\{ \frac{1}{2} \Pi^2 + \frac{1}{2} (\nabla \phi)^2 + m^2 \phi^2 \right\}$

cf. $\frac{1}{2} \frac{p^2}{2m} + \frac{1}{2} \omega^2 x^2$

Other Limitations on Field Theories

1. Locality



Symmetry often allows terms such as $\int d^4x \int d^4x' (\phi(x))^2 V(x-x') (\phi(x'))^2$

However this is action at a distance for general V

- NOT $\mathcal{L}$ for fundamental physics (restrict V)
- $\mathcal{L}$ for low energy / velocity approximation e.g. condensed matter physics.

2. Renormalisation

We can only do calculations with polynomials of fields (usually) up to limited power

- ϕ^2, ϕ^4, ϕ^6 terms, $\mathcal{L}$ depending on space-time dimension
MAX for 3+1, NOT ok in 3+1, MAX in 2+1

- $\sin(\phi^2)$ NOT helpful

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3. Hermitian

To ensure real probabilities need real action & $I = I^*$

Here simple as $\phi(x) \in \mathbb{R} \Rightarrow I \in \mathbb{R}$
 $x^\mu \in \mathbb{R}$

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Complex action $\Rightarrow$ instabilities, non. equilibrium

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