

The new vision on high
redshift clusters with **XMM**:
clusters scaling laws
at **Z** \sim 0.5

Clusters:

<u>Optical light</u>	⇒ Stellar masses + metals	
	⇒ velocity dispersion	⇒ total mass
	⇒ lensing	⇒ total mass
<u>X-ray light</u>	⇒ imagery	⇒ gas mass
	⇒ spectroscopy	⇒ total mass + metals
<u>SZ</u>	⇒ pressure	⇒ gas mass

.

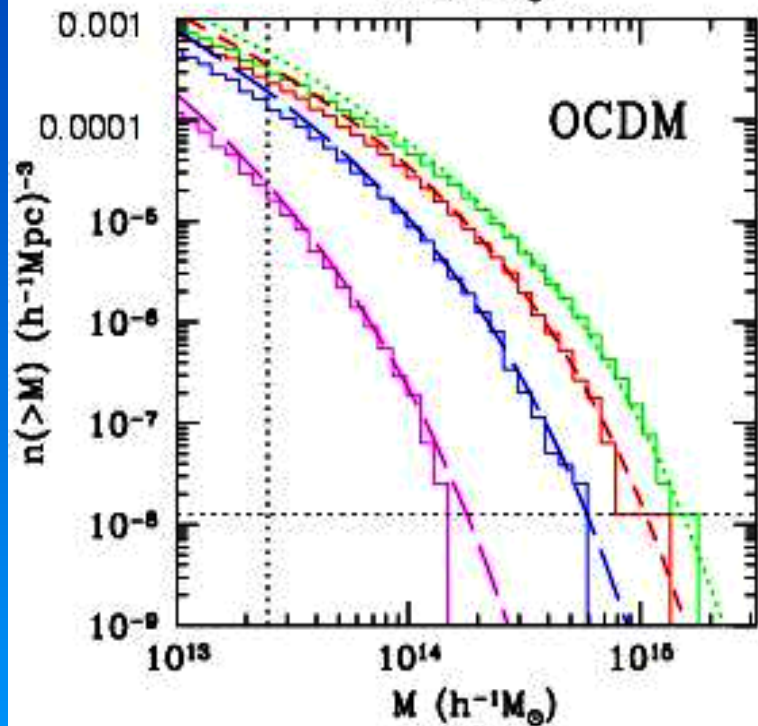
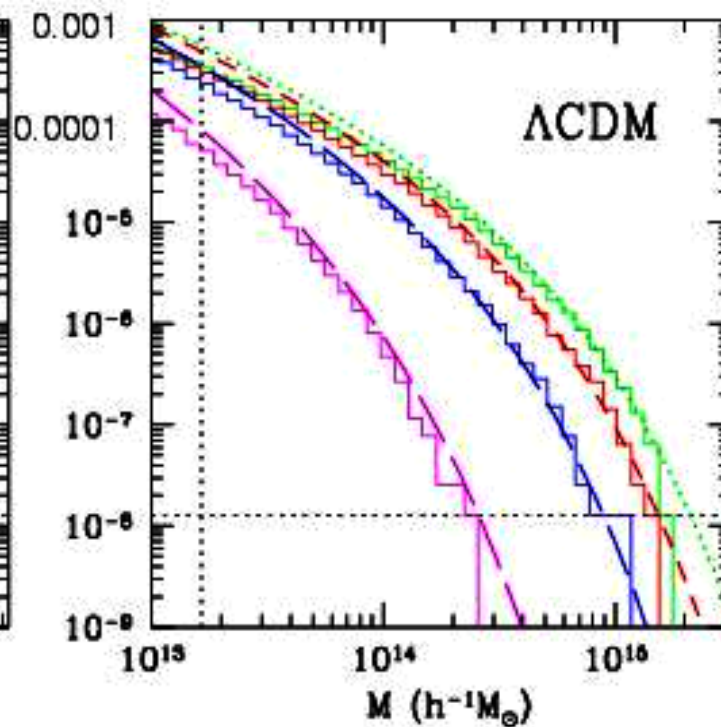
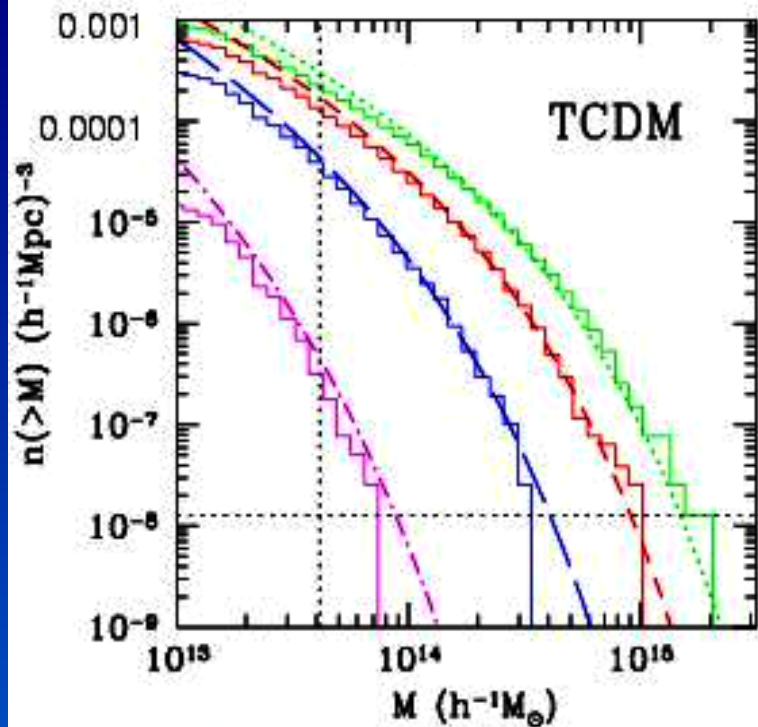
$$2-10 \text{ keV} \Leftrightarrow \mathbf{M} \sim 10^{15} M_{\odot} \Leftrightarrow R \sim 5-20 h^{-1} \text{Mpc}$$

You can (potentially) learn

a lot

in cosmology from

Clusters



$$L_{\text{box}} = 250 \text{ h}^{-1} \text{ Mpc}$$

$$N_{\text{part}} = N_{\text{gr}} = 128^3$$

- $z=0$
- $z=0.21$
- .-.-.- $z=0.55$
- $z=1.40$

Scaling argument for Clusters:

Clusters are geometrically identical

With virial radius-mass relation

$$M = \frac{4\pi}{3} \rho_0 (1+z)^3 (1+\Delta) R_v^3$$

i.e.

$$R_v = \sqrt[3]{\frac{3M}{4\pi\rho_0(1+\Delta)(1+z)}}$$

Mass-Temperature Relation :

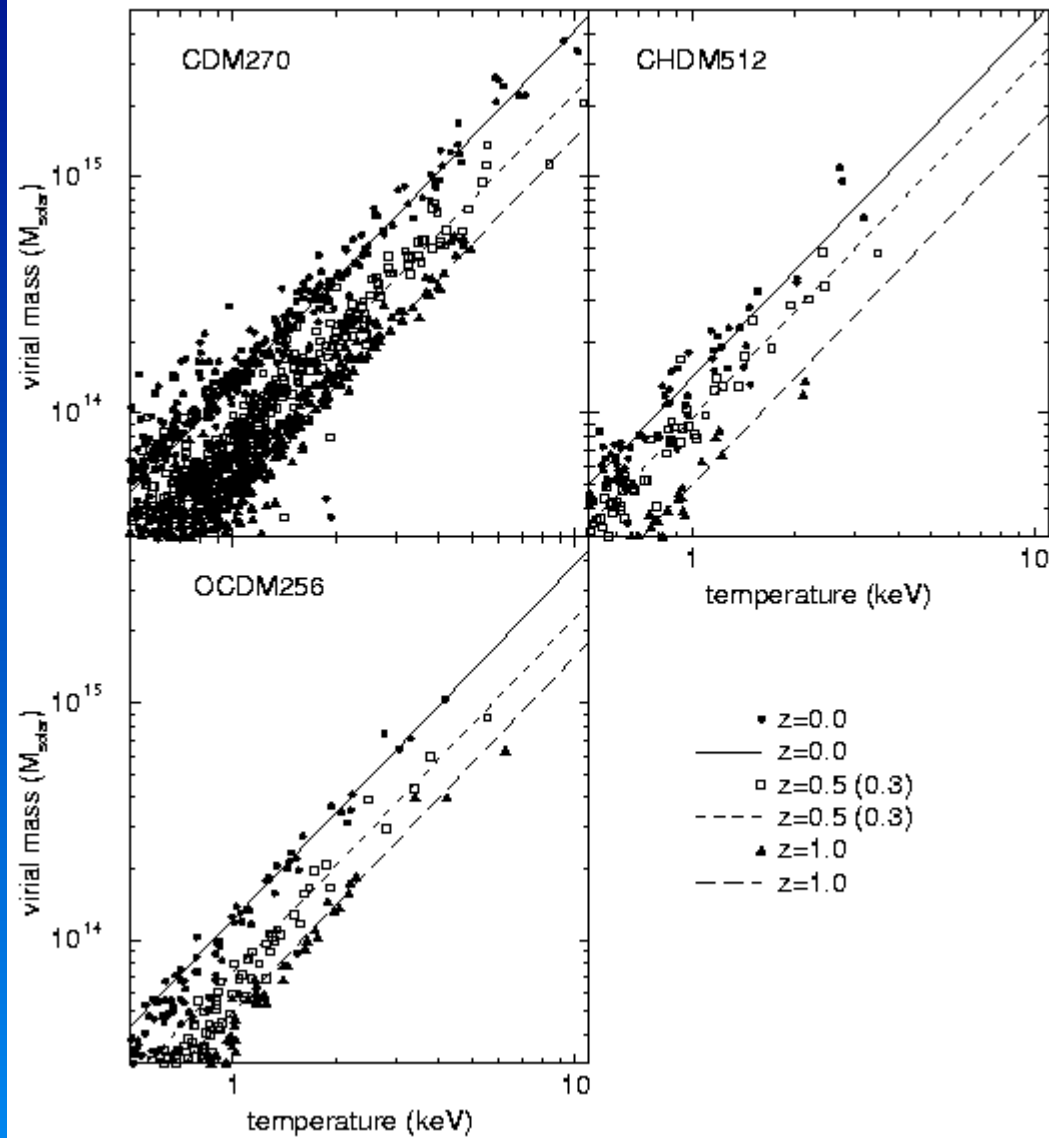
$$T \propto GM/r$$

whatever you do with gravity...

$$T_x \simeq A M^{2/3} (\Omega \Delta)^{1/3} (1+z) \text{ keV}$$

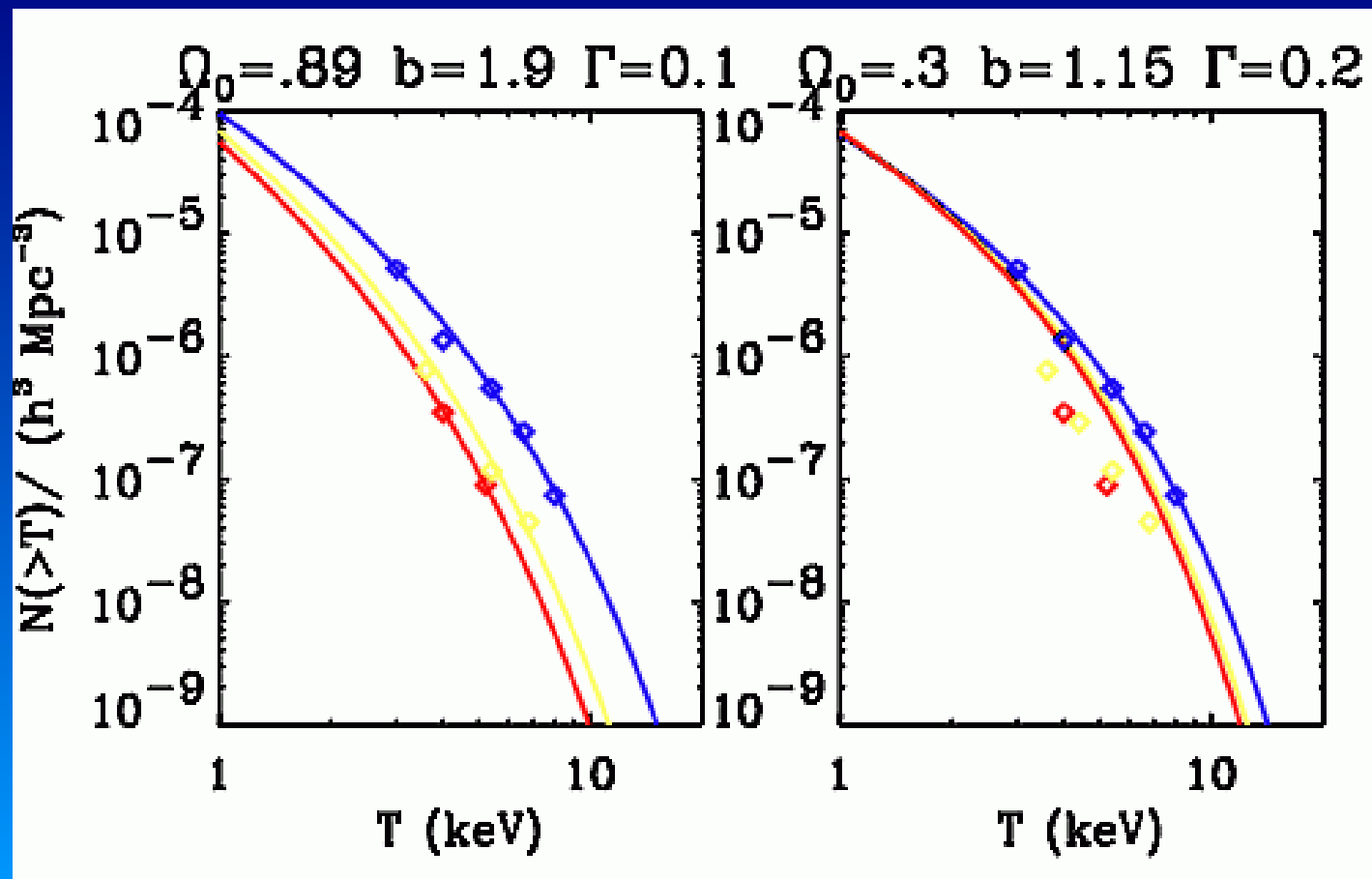
Numerical simulations, Bryan & Norman, 1998

STATISTICAL PROPERTIES OF X-RAY CLUSTERS



Bryan & Norman (ApJ 495 80 1998)

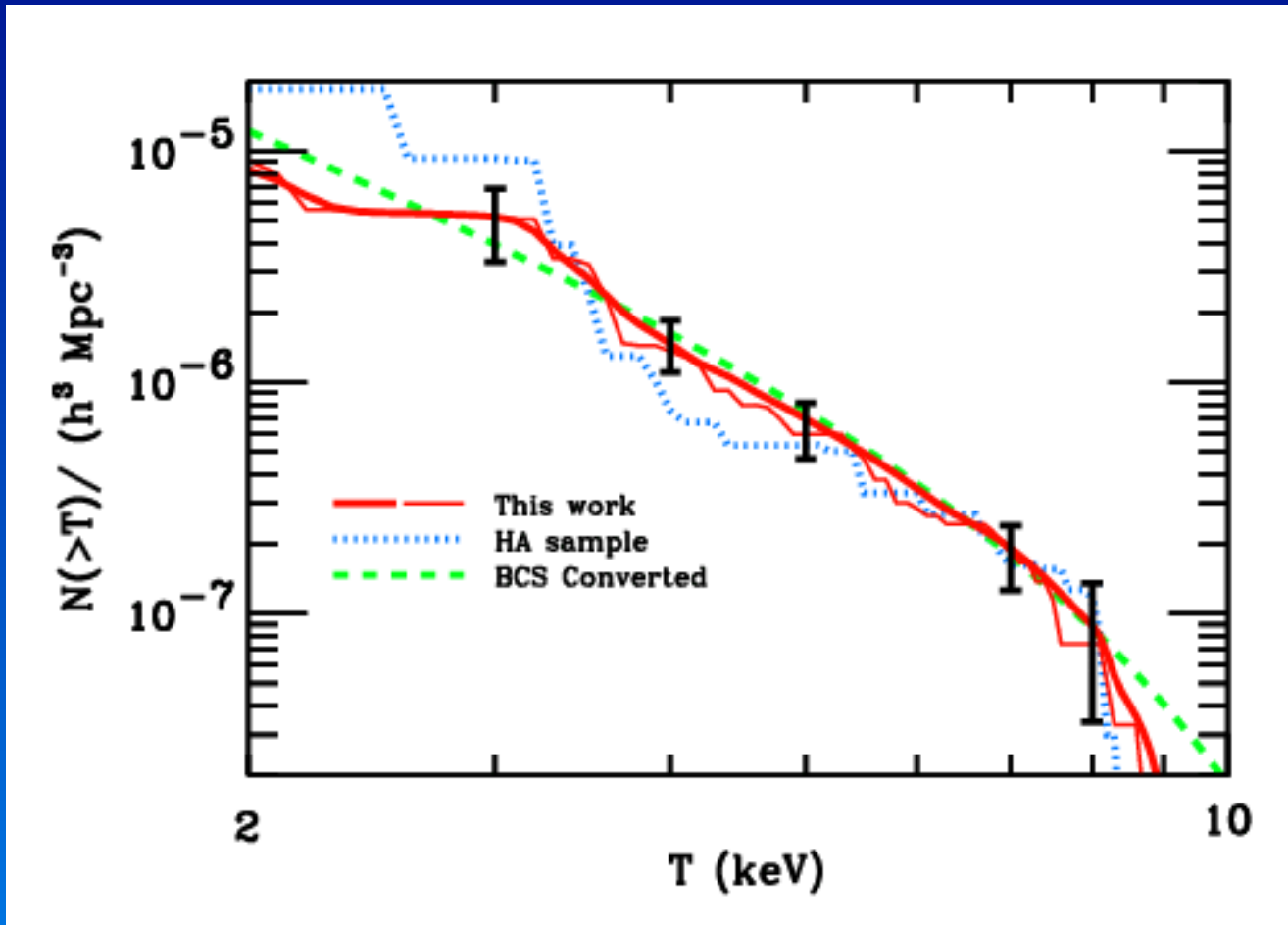
Principle



X-ray clusters allow
precision
cosmology...

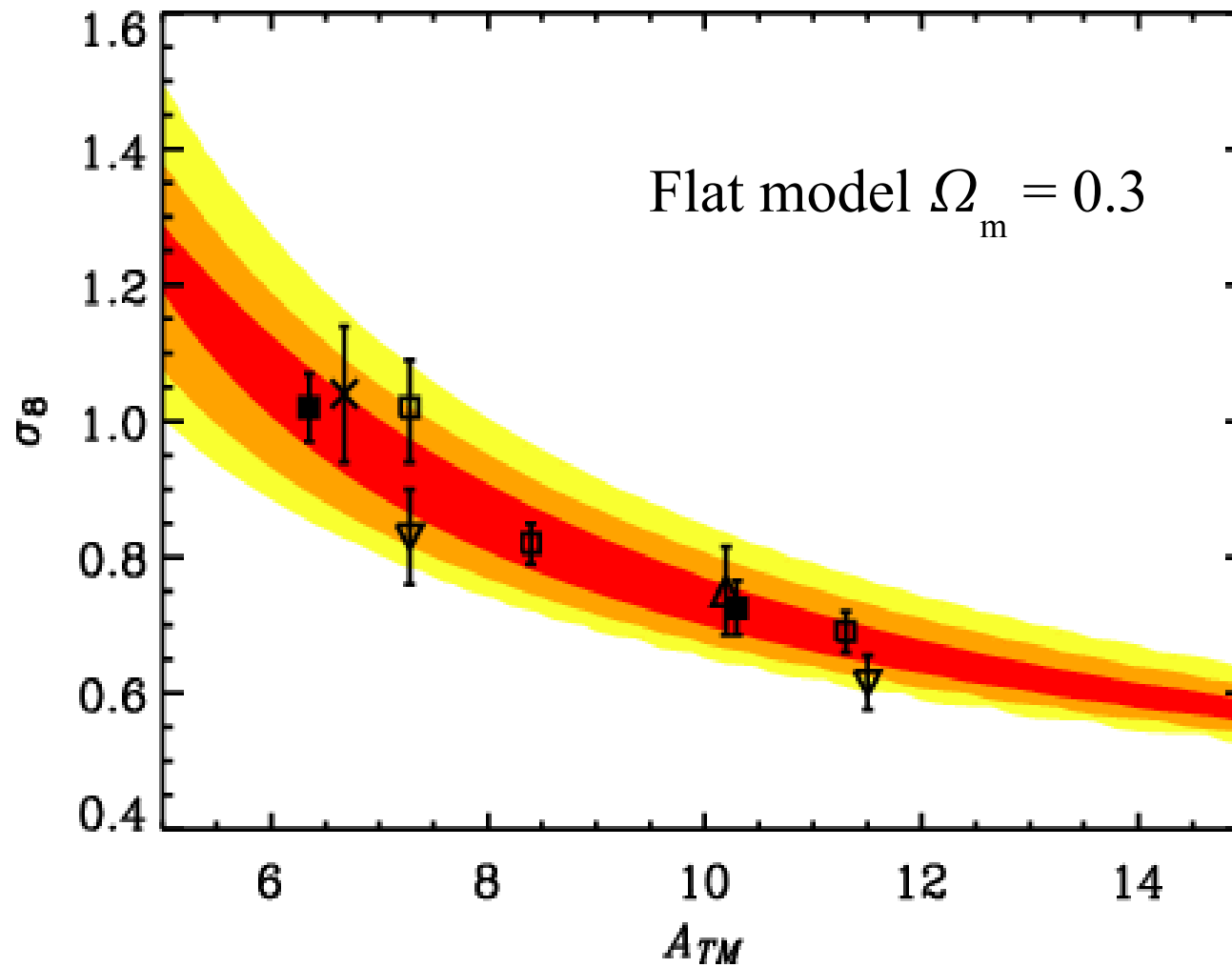
σ_8

Estimated $N(T)$ at $z = 0.05$



50-60 ROSAT clusters \Rightarrow “Convergence” : (Markevitch, 1998), Blanchard et al. (2000), Pierpaoli et al (2001), Ikebe et al (2002), Pierpaoli et al (2002)

σ_8 from X-ray clusters:



$$\Omega_m$$

From X-ray Clusters

Number evolution

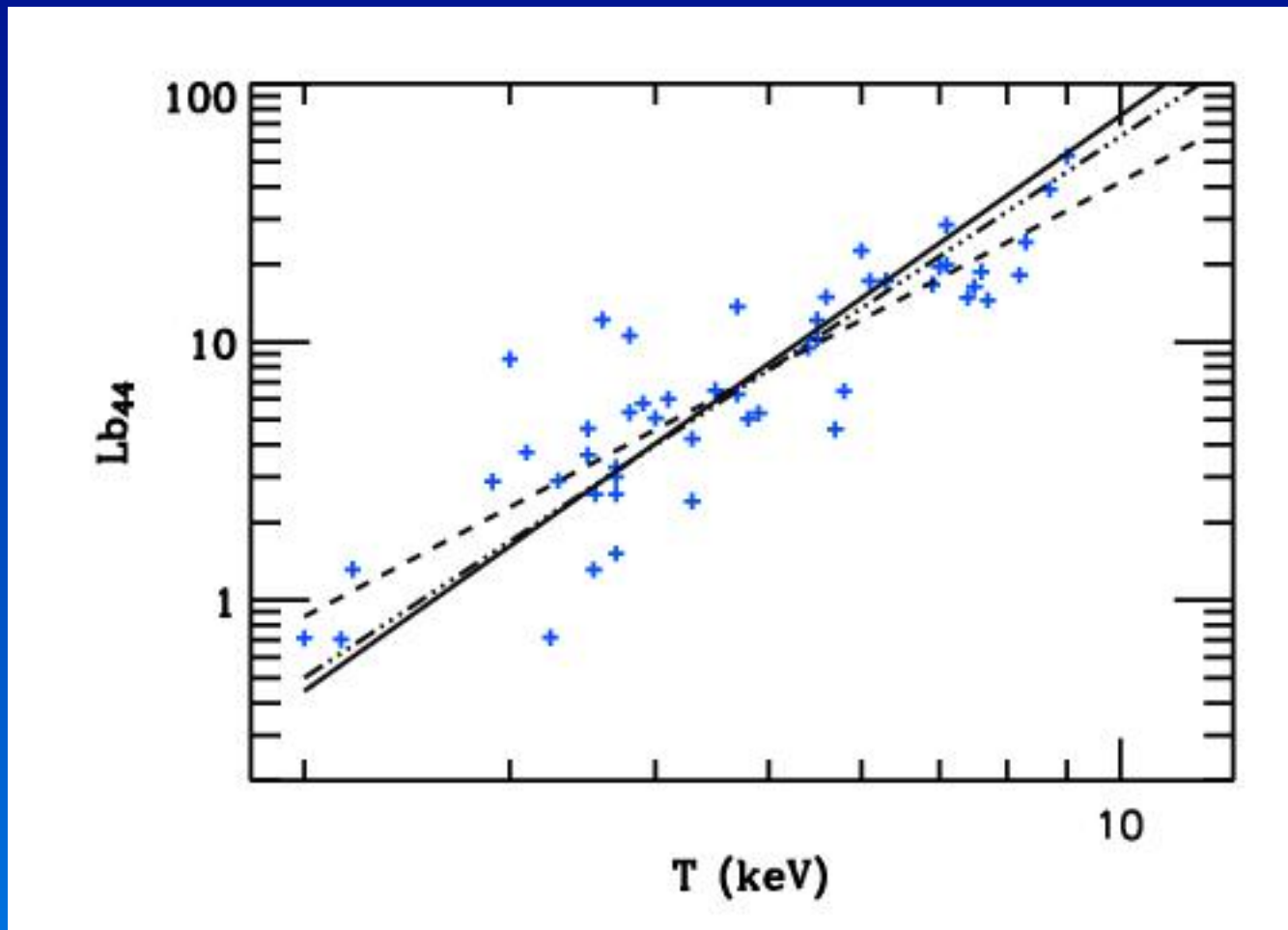
Mass-Luminosity Relation :

$$L_x \propto n^2 T^{1/2} V$$

$$L_x \simeq B M^{4/3} \Omega^{1/6} \Delta^{7/3} (1+z)^{3.5}$$


$$L_x \propto T^2 \Delta^{1/2} (1+z)^{1.5}$$

Observed Temperature -Luminosity Relation



$$L_x \propto T_x^{\sim 3}$$

No standard scaling for L-T
and its evolution...

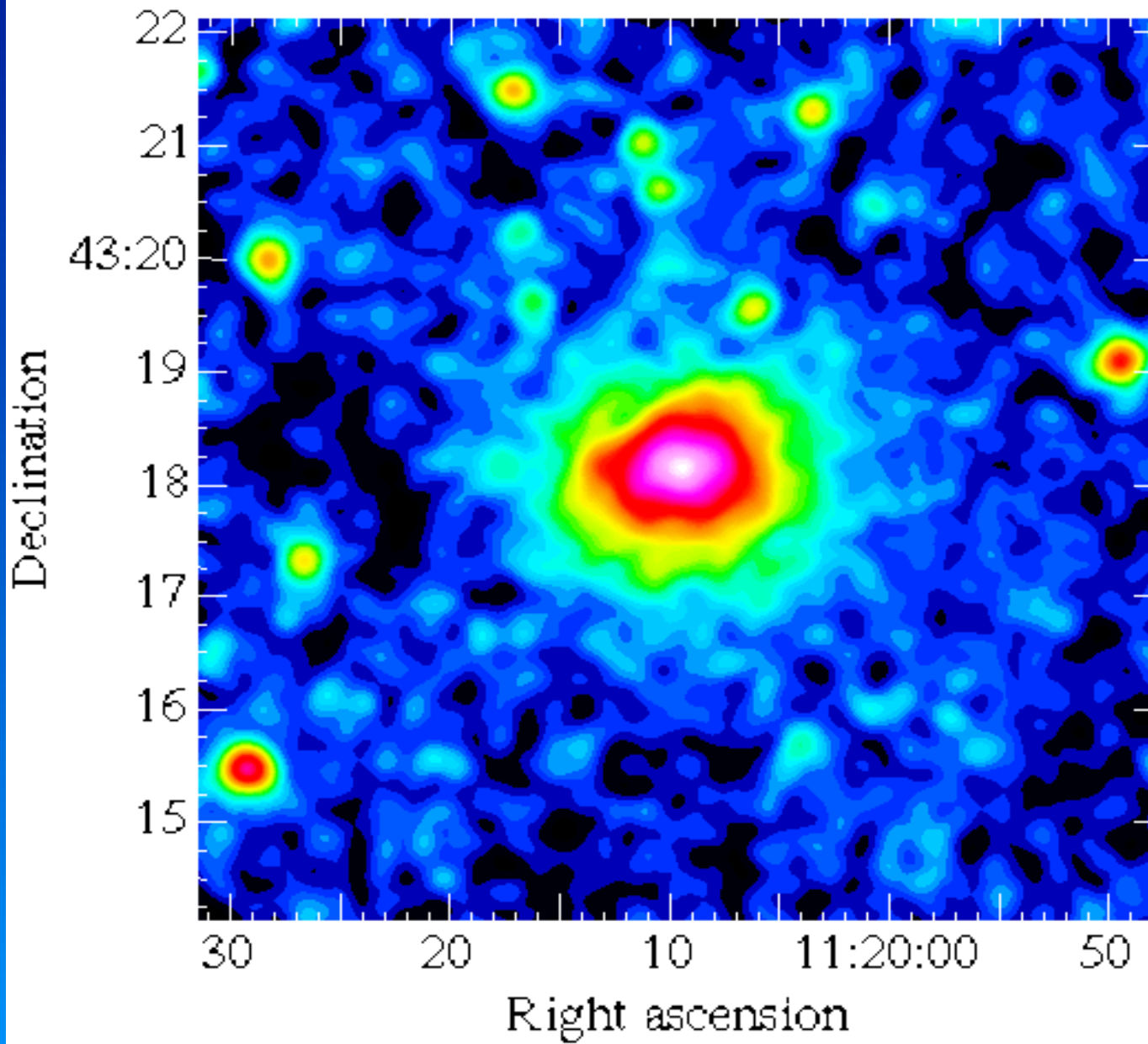
XMM

Ω -Project

**X-ray properties of distant SHARC clusters
for Cosmology
with a complete flux limited survey.**

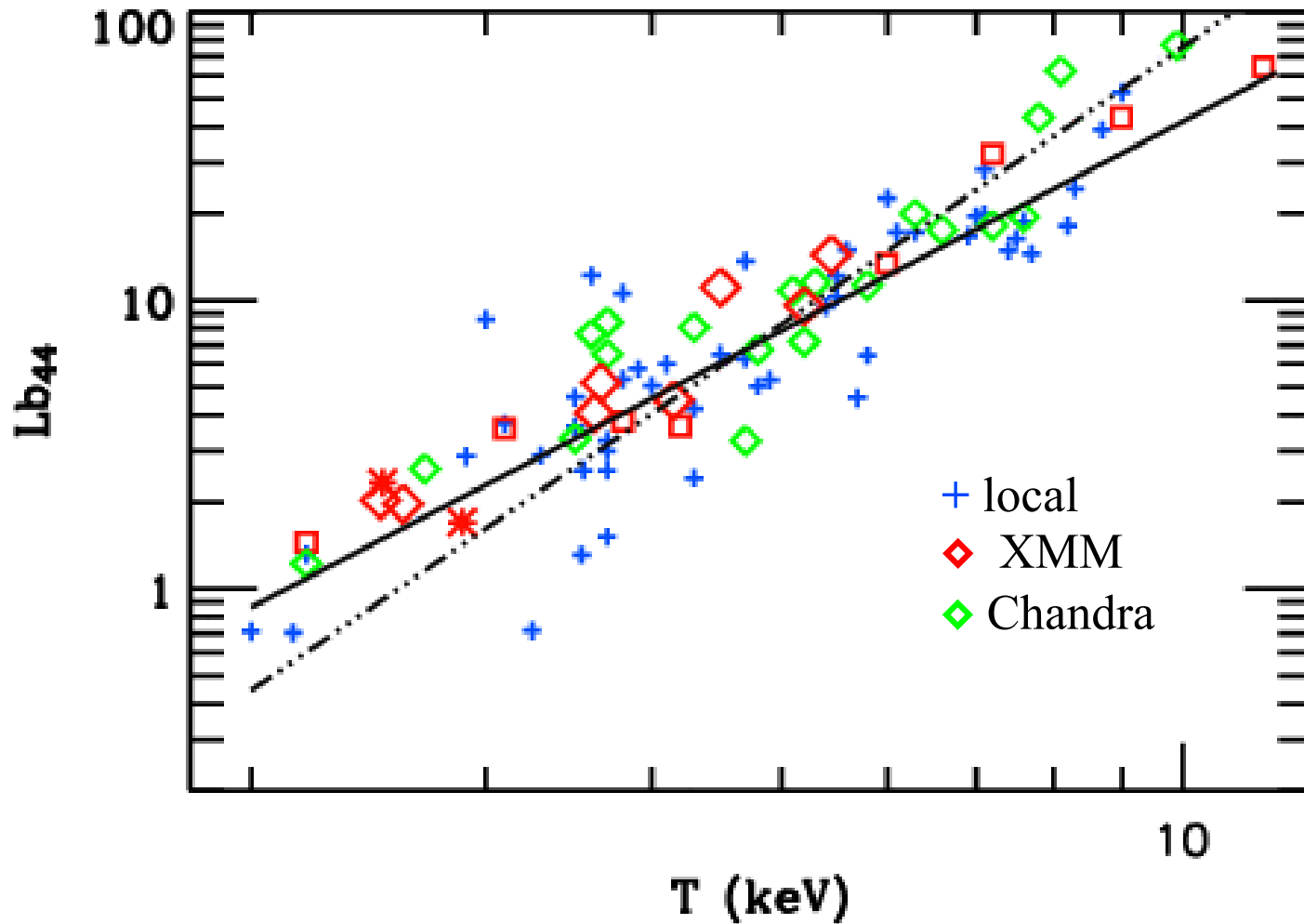
$Z \approx 0.6$

X
M
M



R
X
J
1
1
2
0

XMM Lx-Tx evolution



Conclusion on evolution:

❖ remarkable convergence

$$\left(\frac{L_x}{T_x}\right)_z = \left(\frac{L_x}{T_x}\right)_{z=0} (1+z)^\beta$$

with $\beta = 1.52 \quad 0.28$ D.Lumb et al., 2003

in full agreement with ASCA (Sadat et al., 1998; Novicki et al., 2003....), Chandra (Vikhlinin et al, 2002), and more recent XMM analyses (Kotov & Vikhlinin, 2006; Maughan et al. 2006)

Number counts:

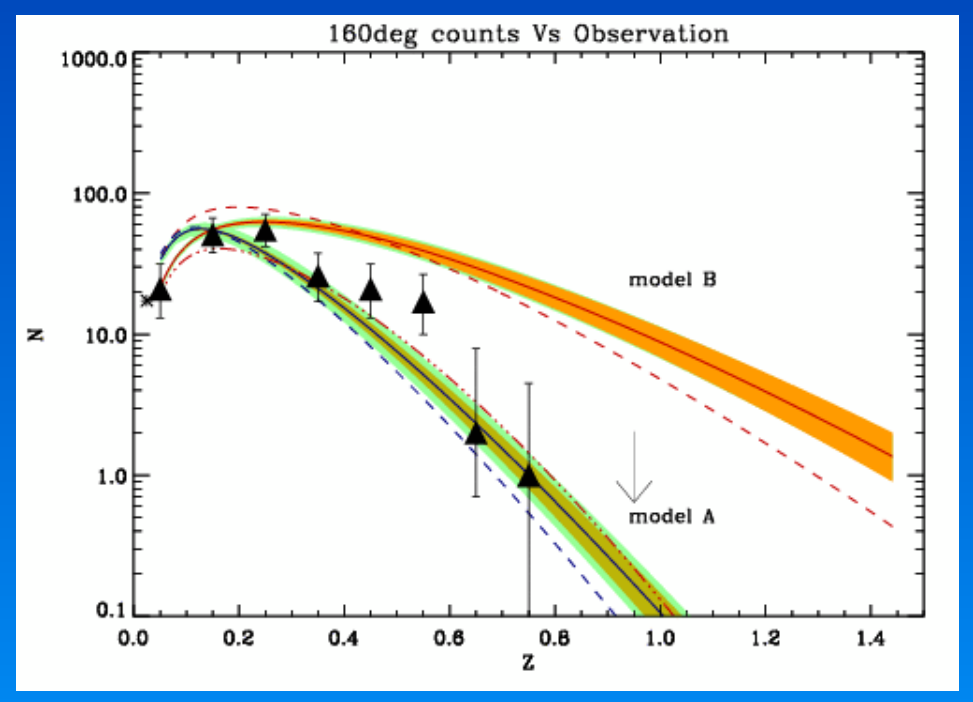
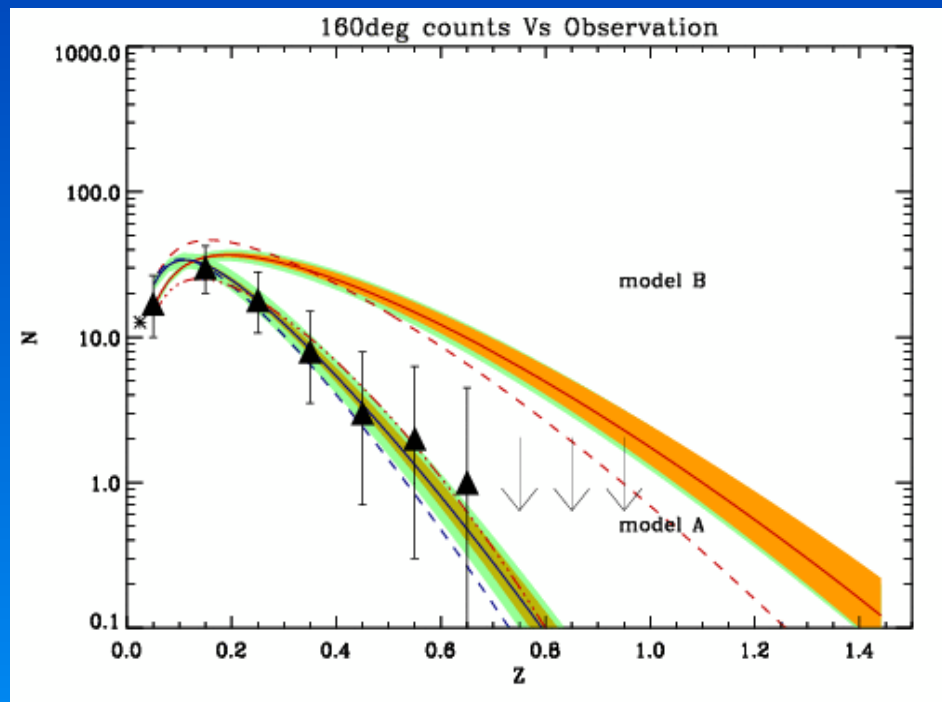
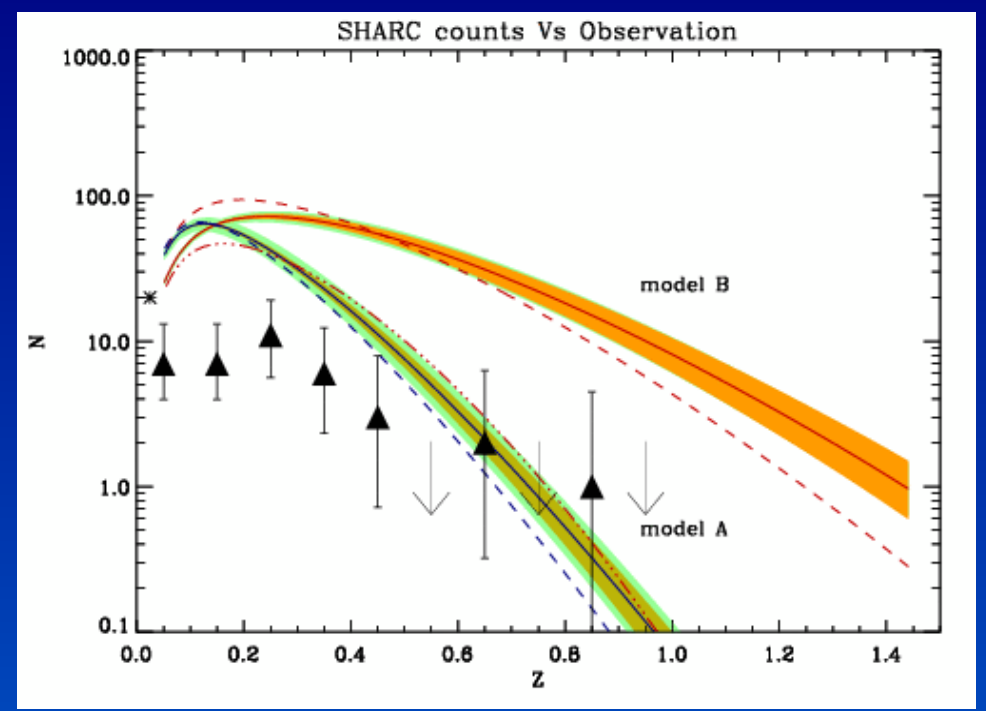
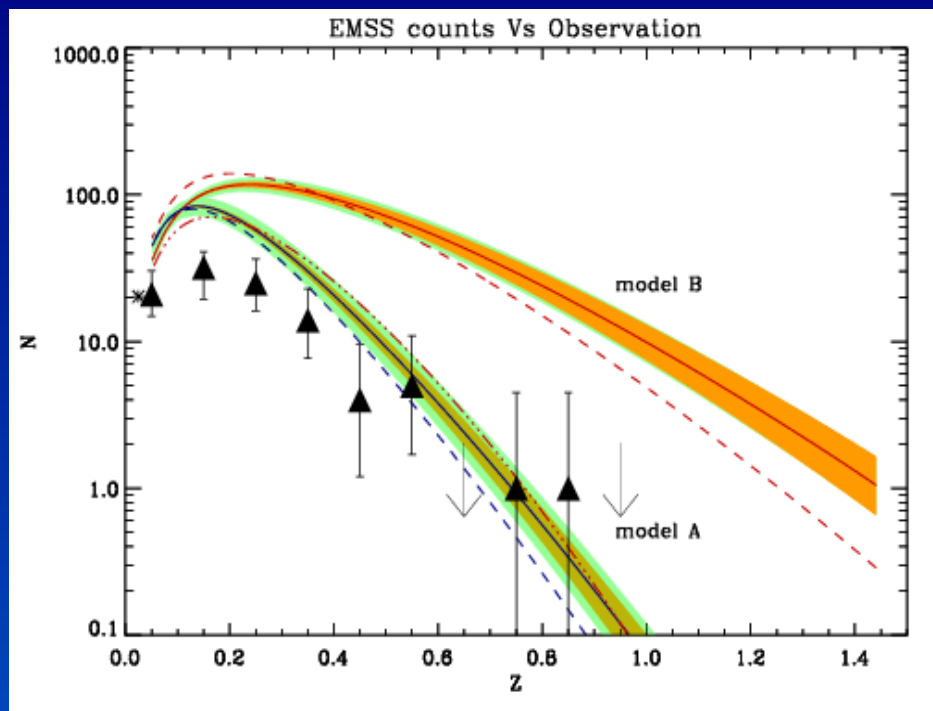
Vauclair et al, 2003
A&A 412, L37

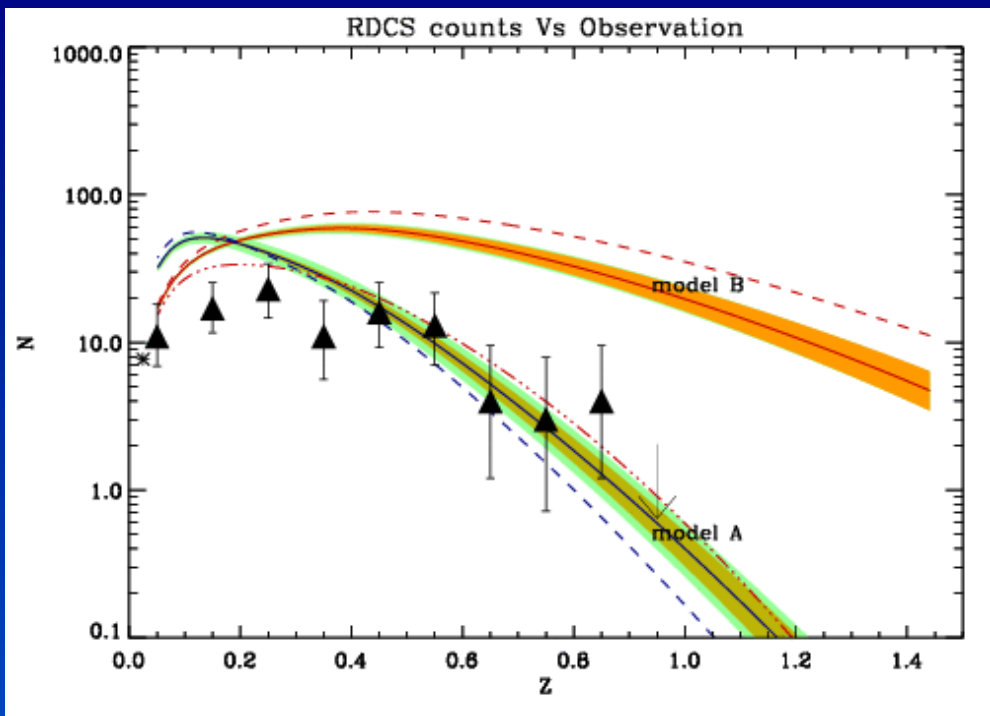
300 clusters
with $z > 0.3$

Method:

$$f_x \rightarrow L_x \rightarrow s, T_x$$

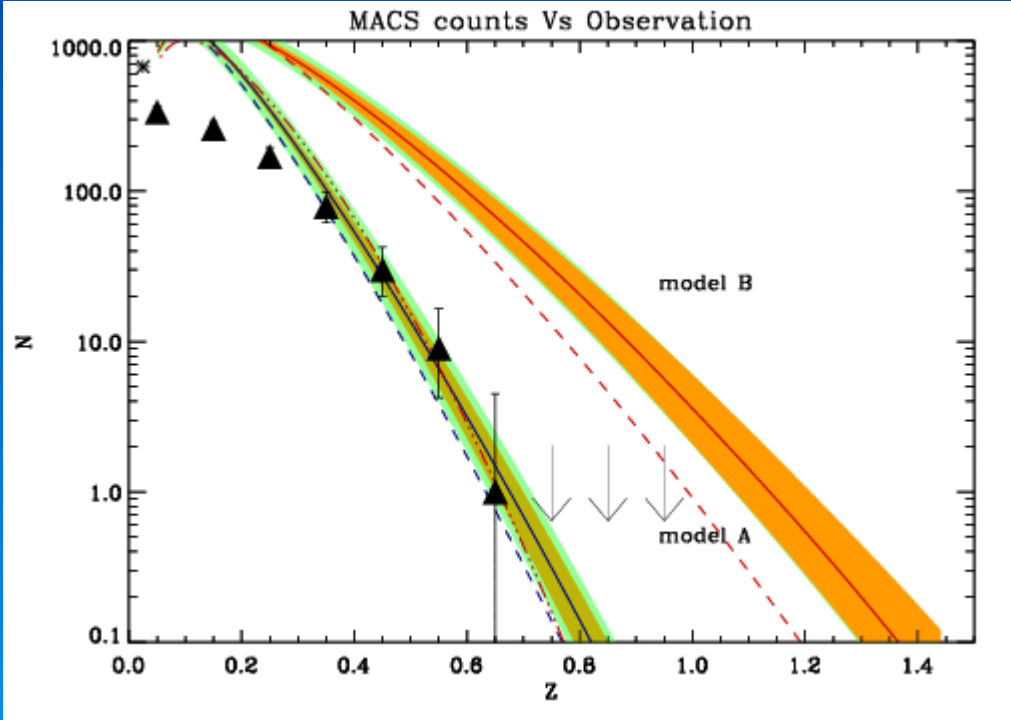
$$\begin{aligned} N(> f_x) &= \int_0^{+\infty} \int_0^{+\infty} s(T, z) N(T, z) dT dV(z) \\ &> \sim \int_0^{+\infty} N(> T(z)) dV(z) \end{aligned}$$



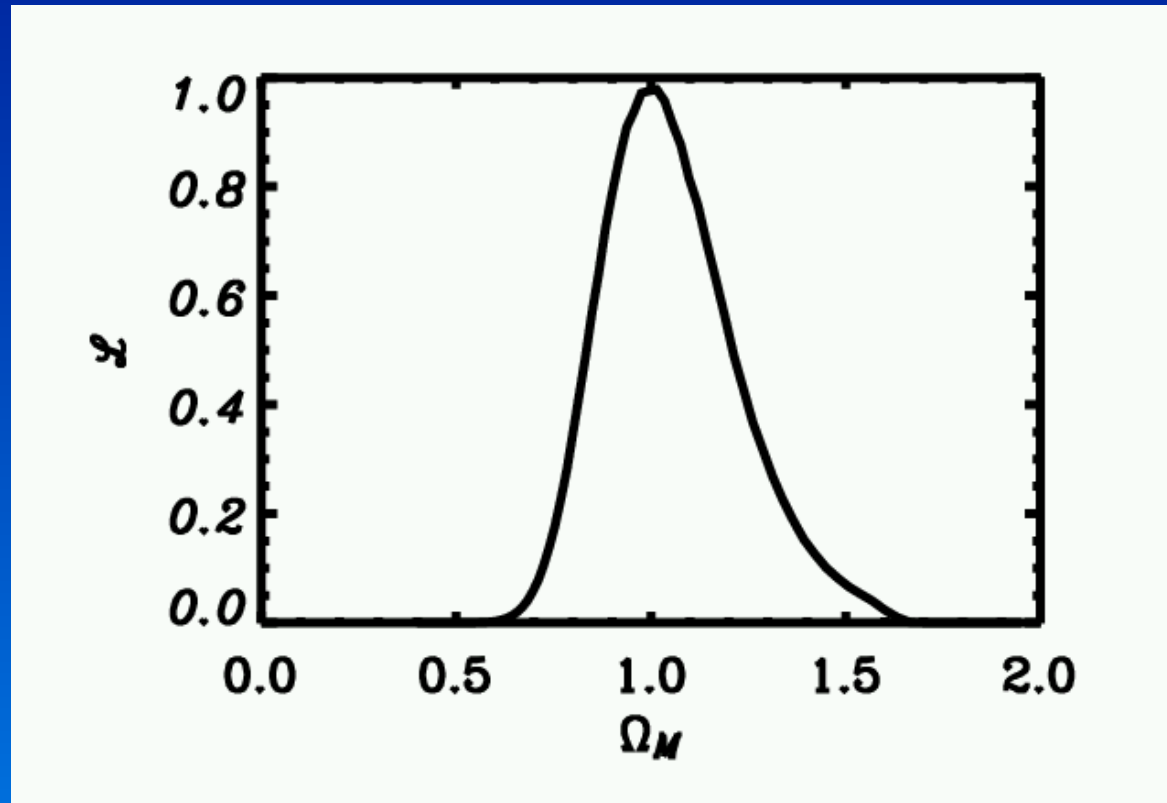


RDCS: 50 deg²
 $f_x \approx 3 \cdot 10^{-14}$ erg/s/cm²

MACS: 22 000 deg²
 $f_x \approx 10^{-12}$ erg/s/cm²



Likelihood analysis:



(Vauclair et al., 2004)

$$\Omega_m = 0.99 \pm 0.15 \pm 0.15$$

Conclusion at that point is :

Clusters observations are
inconsistent with
self similar models in
concordance cosmology!

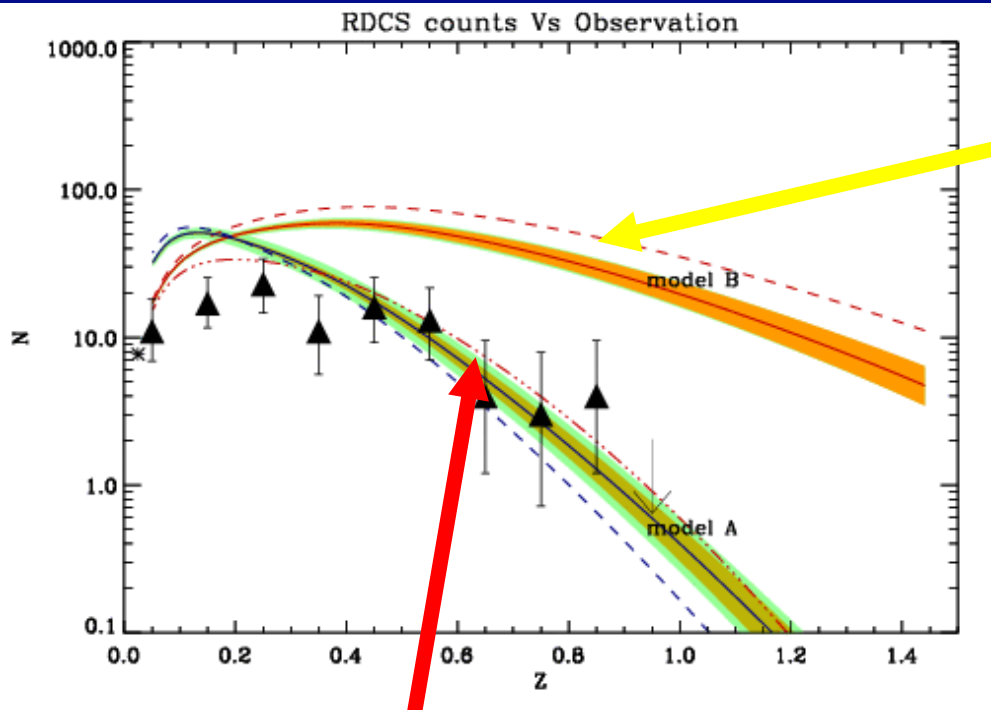
Kill the Mass-Temperature Relation :

$$T \propto GM/r + \dots \propto GM/r / (1+z)$$

i.e. \sim forget gravity...

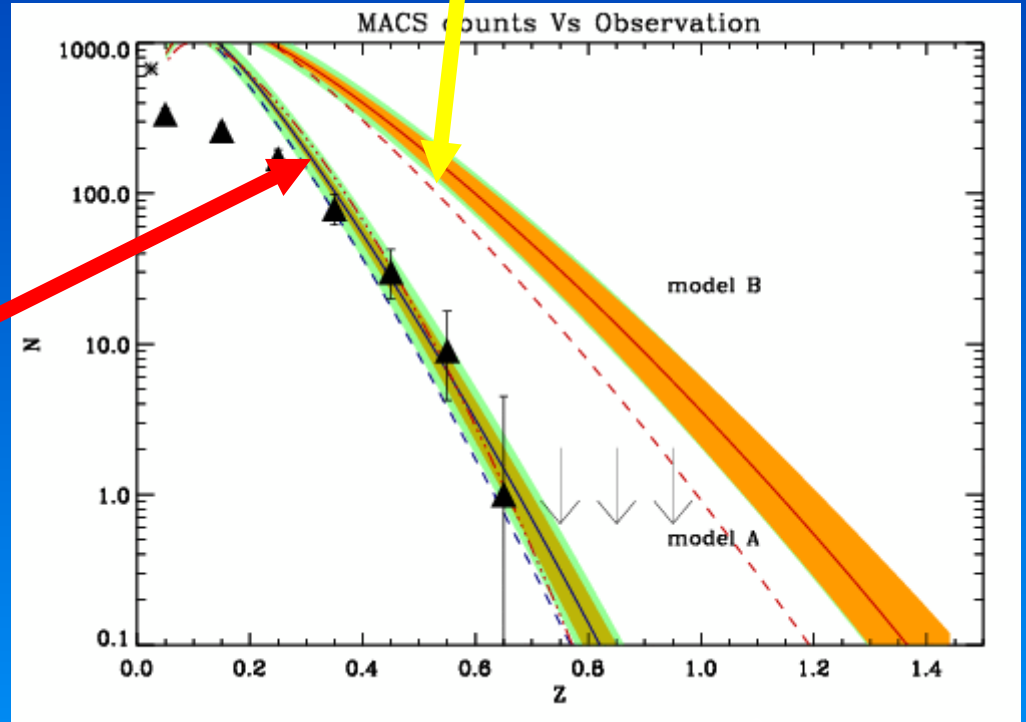

$$T_x \simeq A M^{2/3} (\Omega \Delta)^{1/3} (1+z)^0 \text{ keV}$$

RDCS: 50 deg², $f_x \approx 3 \cdot 10^{-14}$ erg/s/cm²



Concordance model
+
Standard M-T scaling law

Concordance model
+
Revised M-T scaling law



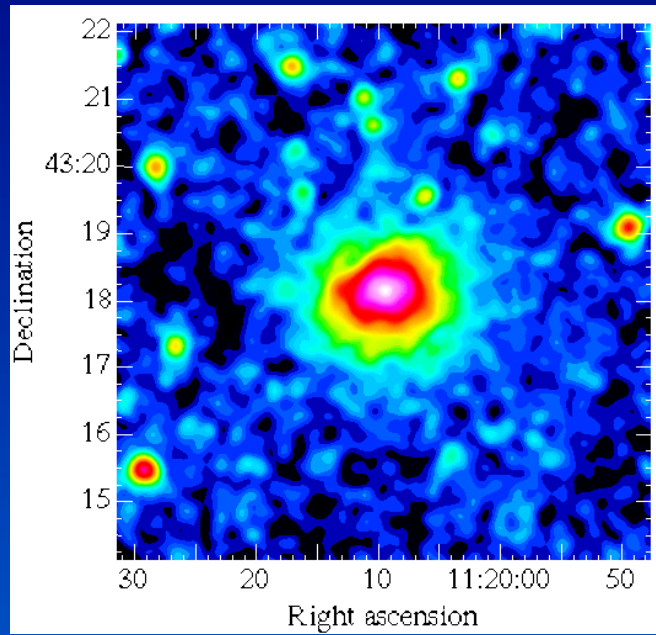
MACS: 22 000 deg² $f_x \approx 10^{-12}$ erg/s/cm²

|| Ω_m From X-ray Clusters

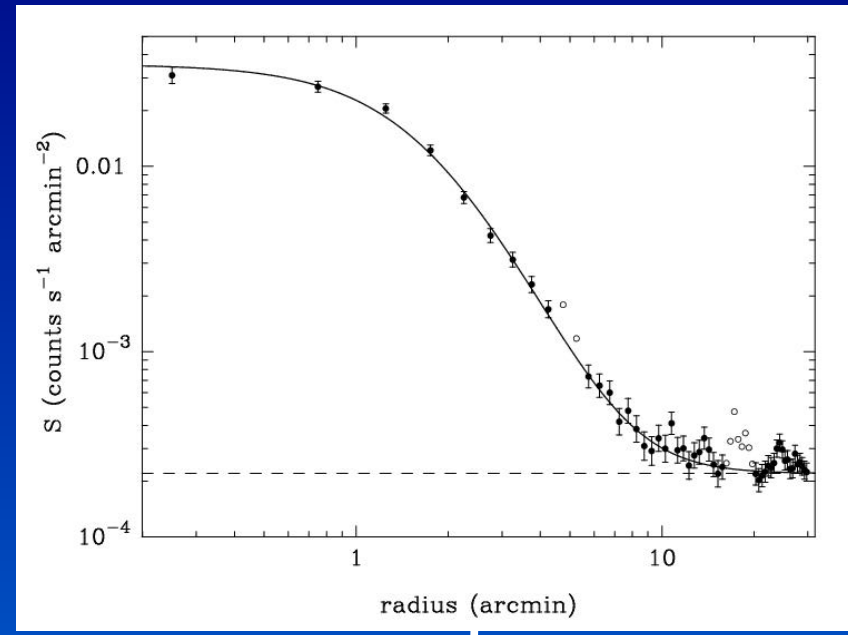
Baryon Fraction evolution
in the XMM Ω -project

(Sadat et al., A&A 2005)

What do you do with a cluster?

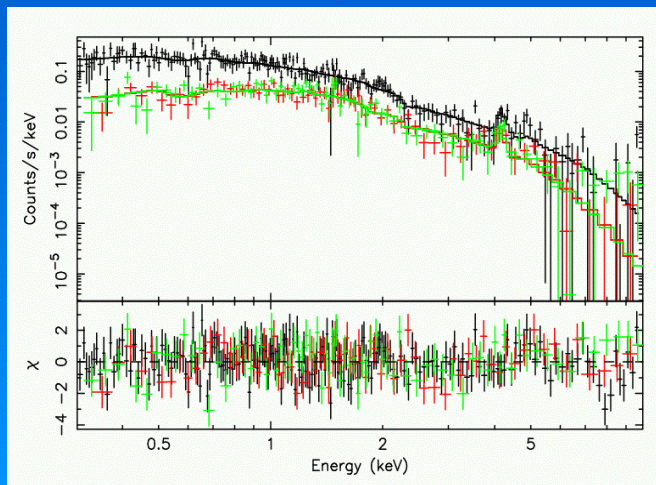


Fit

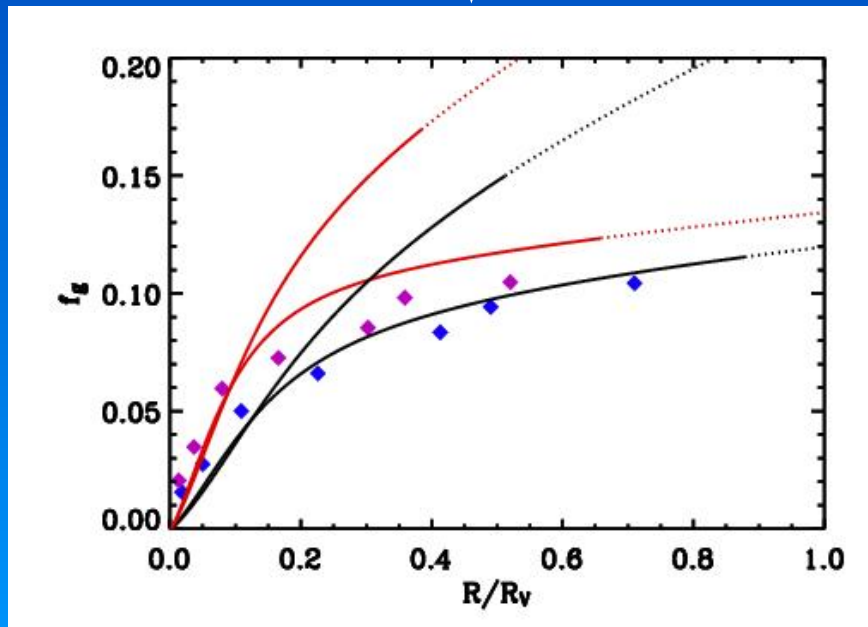


Gas mass $M_g(r)$

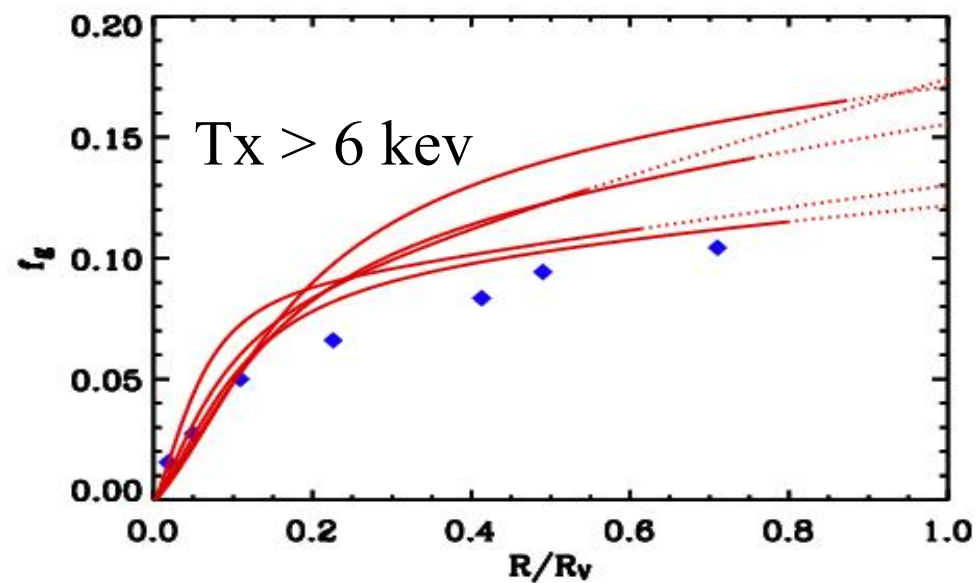
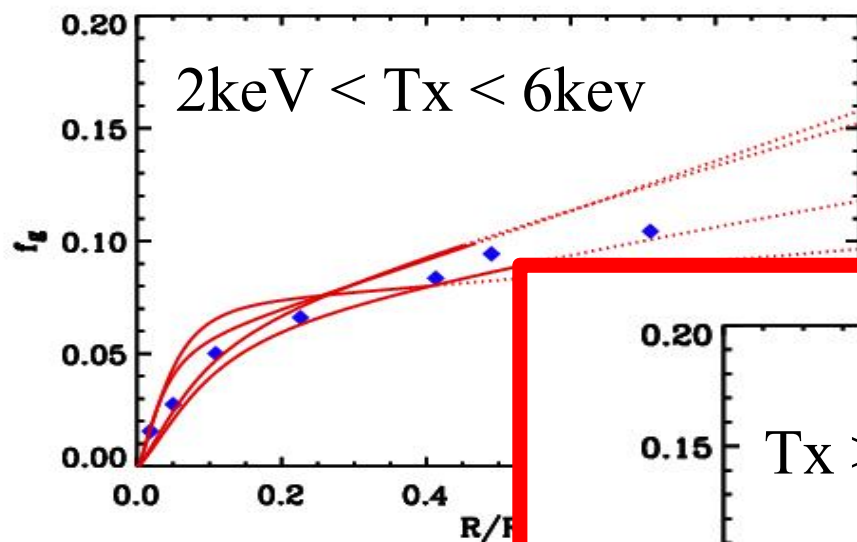
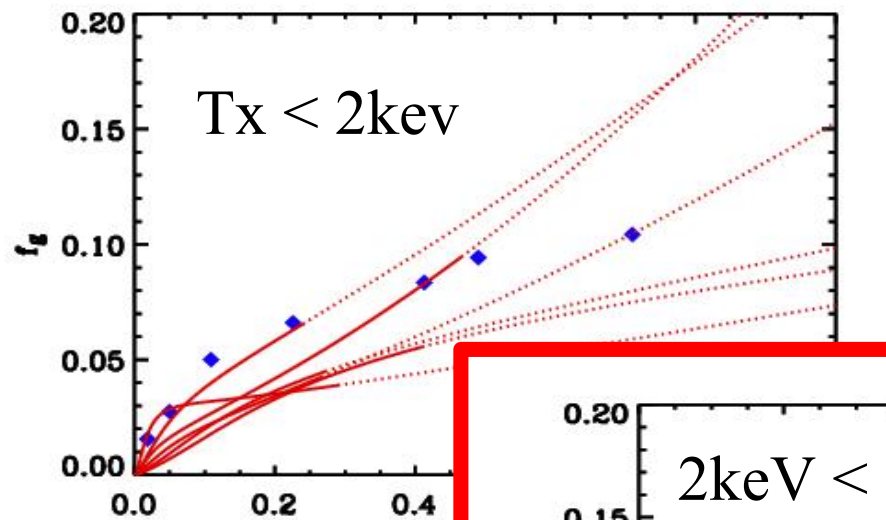
X-ray spectrum



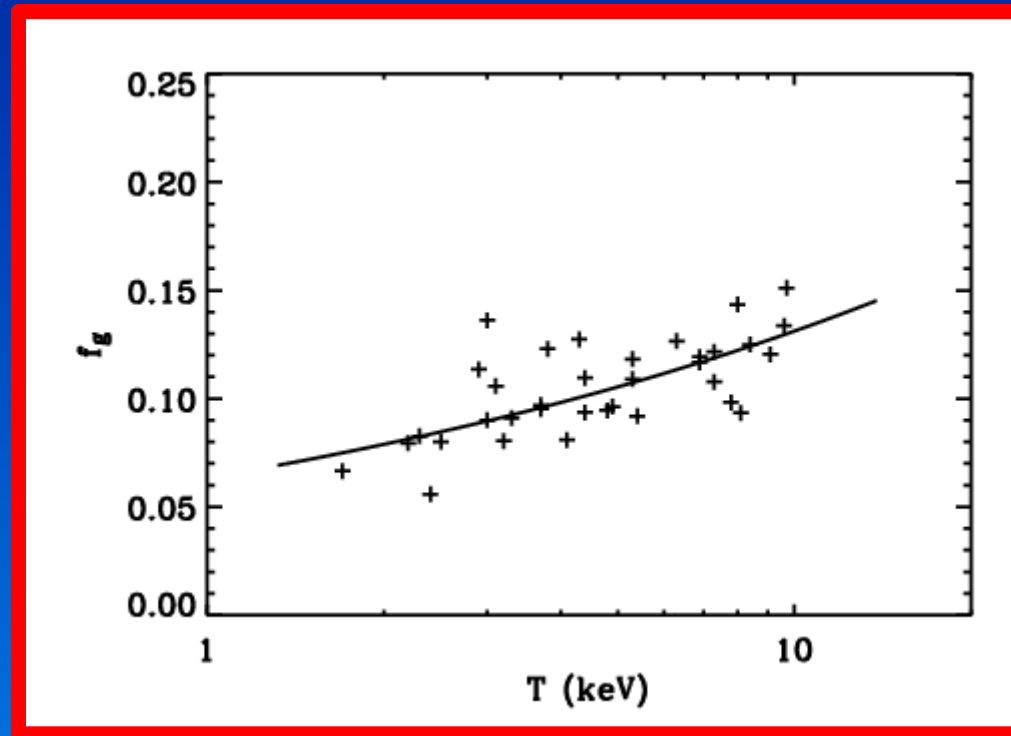
$M(r)$



Barvon Fraction @ $z = 0$

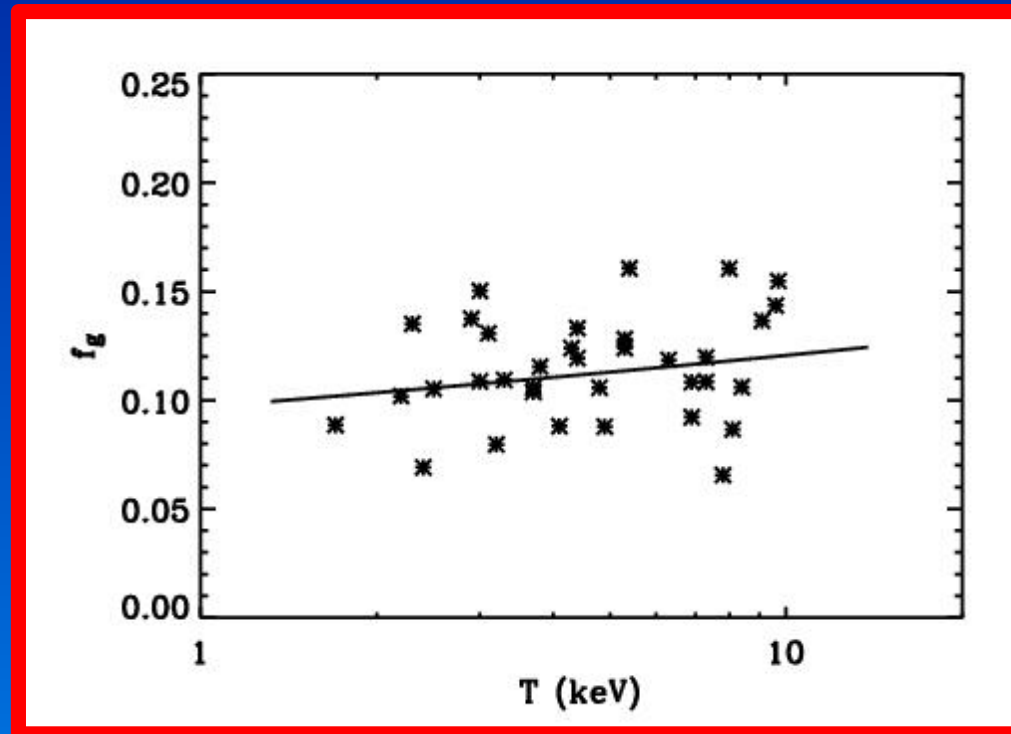


Baryon Fraction @ $z = 0$



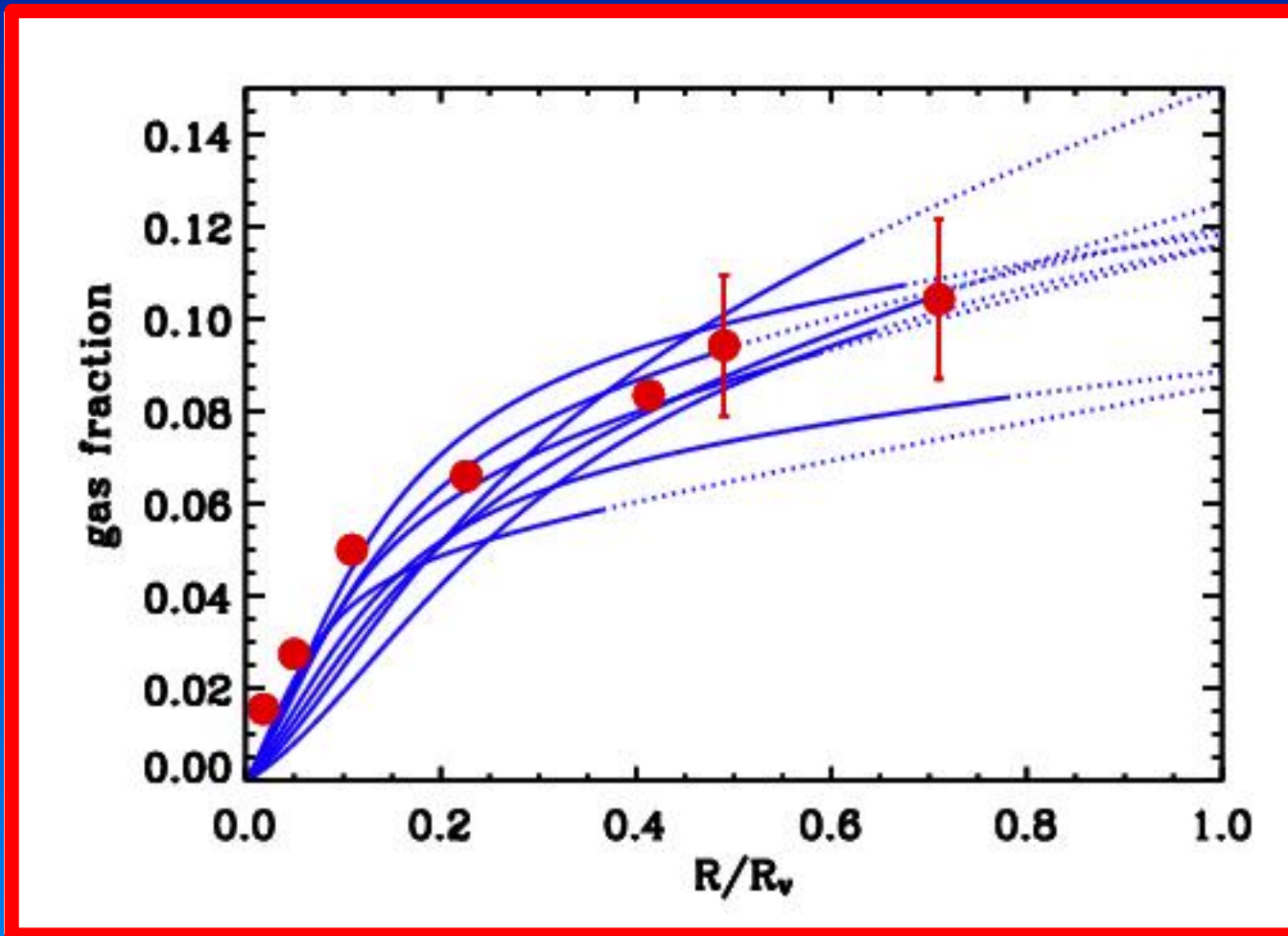
R_{2000} in Vikhlinin, Forman, Jones 1999 ($\sim 35-45\% R_V$)

Baryon Fraction @ $z = 0$

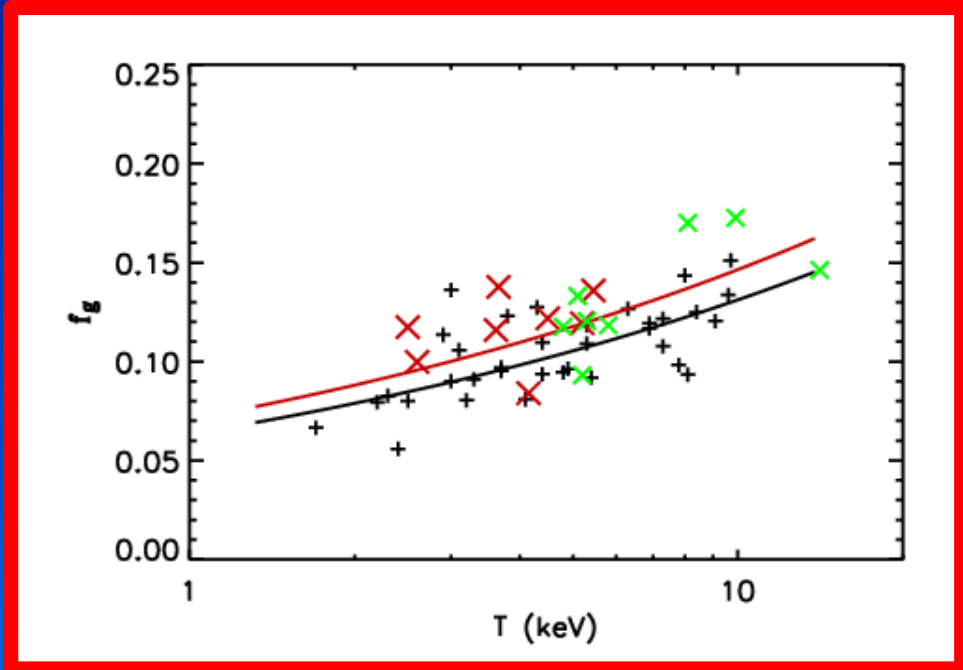


R_v in Vikhlinin, Forman, Jones 1999

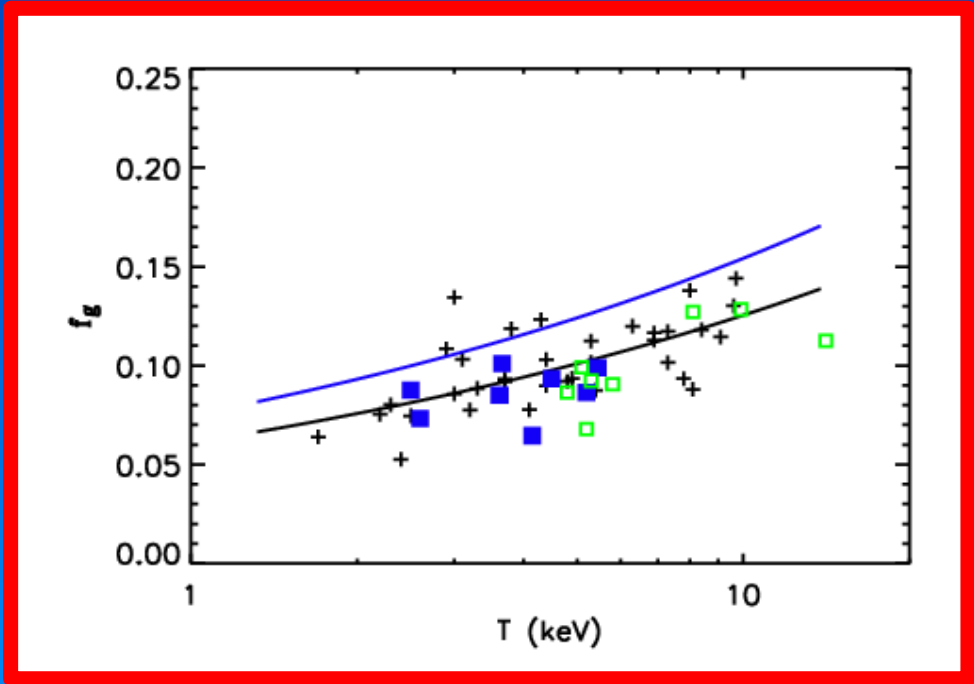
Baryon Fraction @ $z = 0.6$



Baryon Fraction @ $z = 0.6$



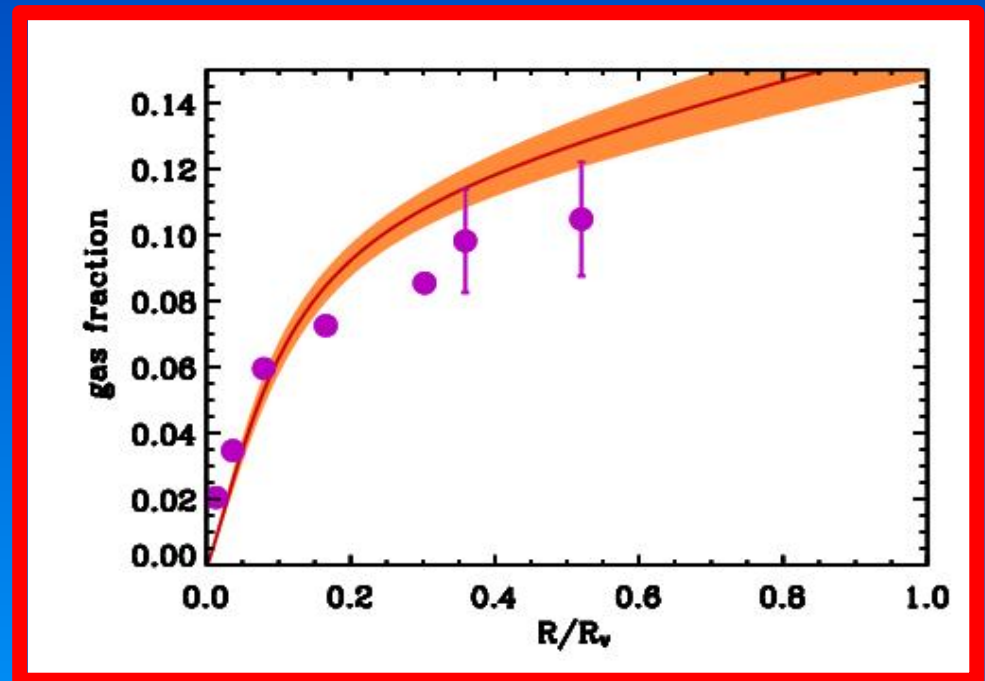
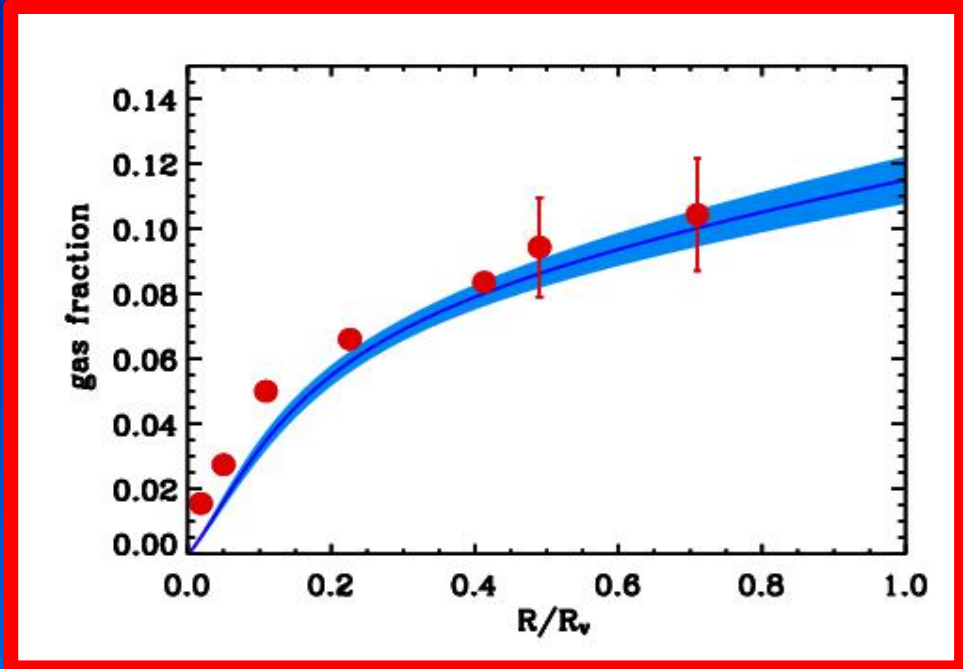
$\Delta = 2000$



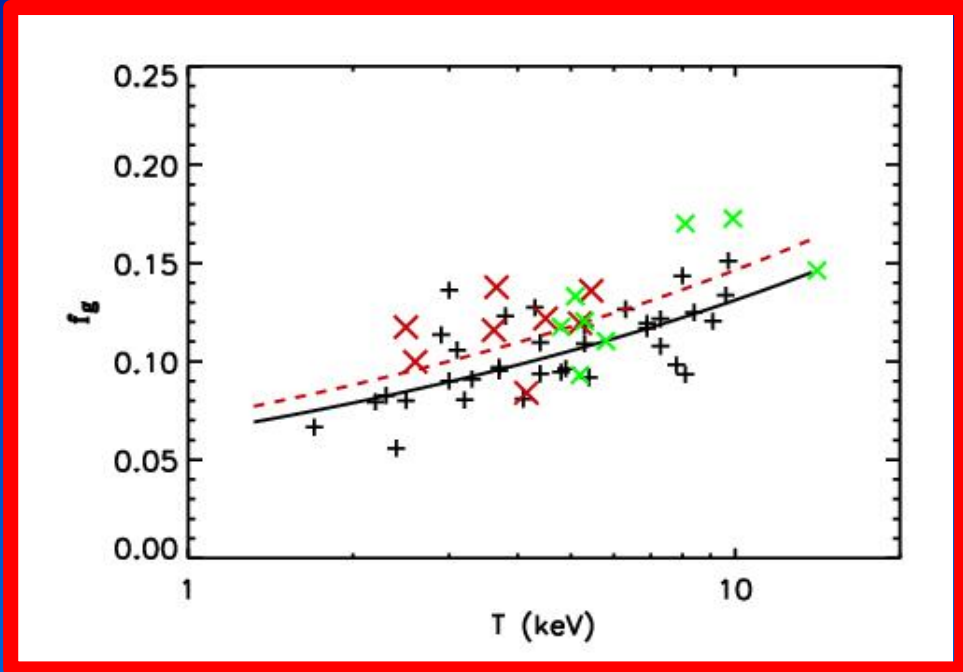
Baryon Fraction @ $z = 0.6$

Internal structure

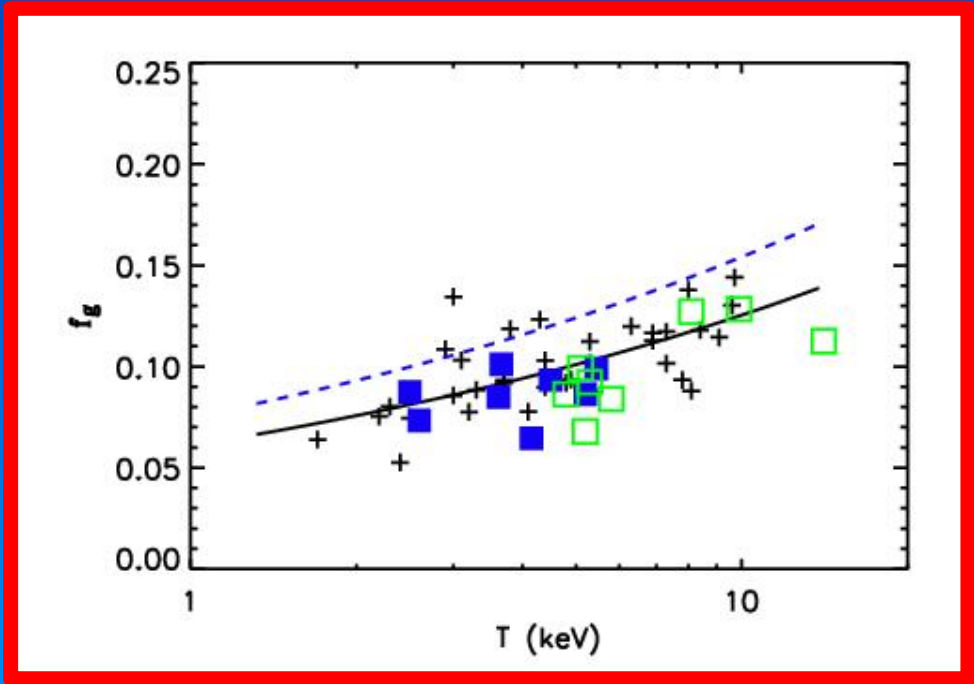
is complex...



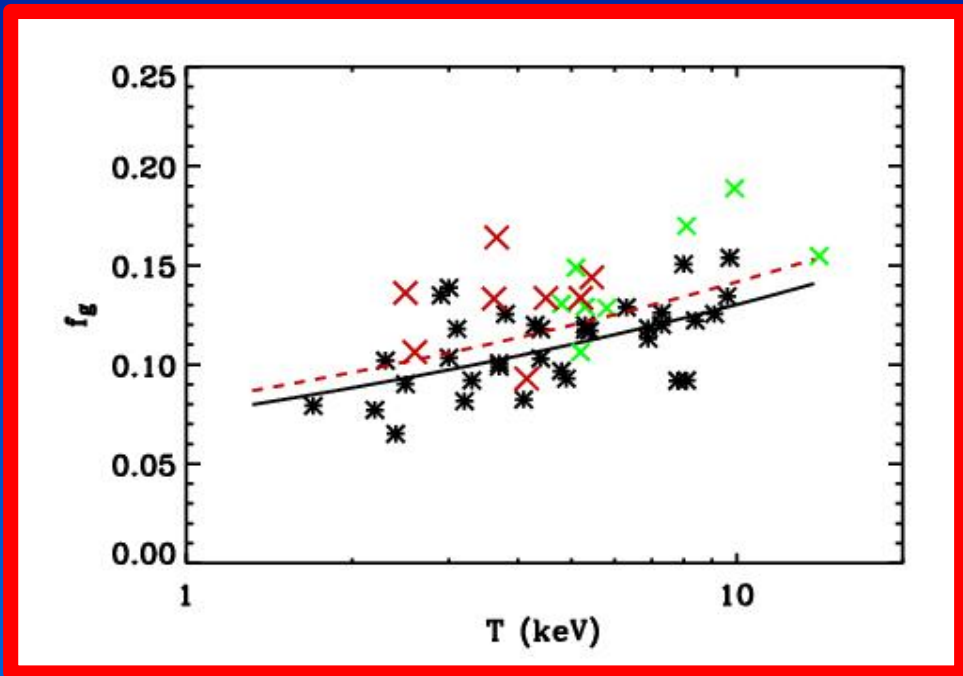
Baryon Fraction @ $z = 0.6$



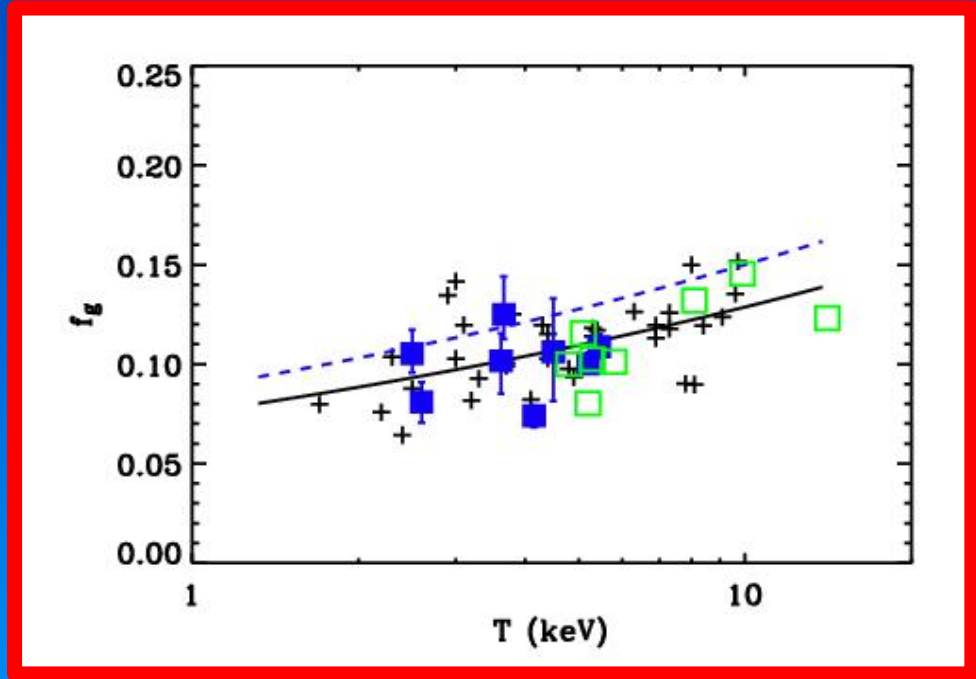
$\Delta = 2000$



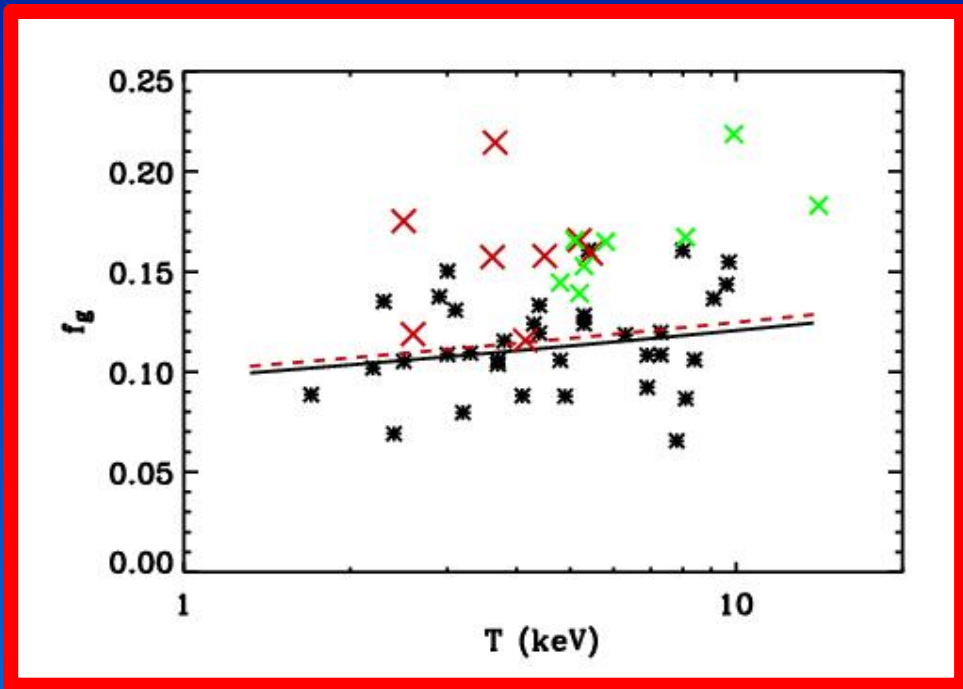
Baryon Fraction @ $z = 0.6$



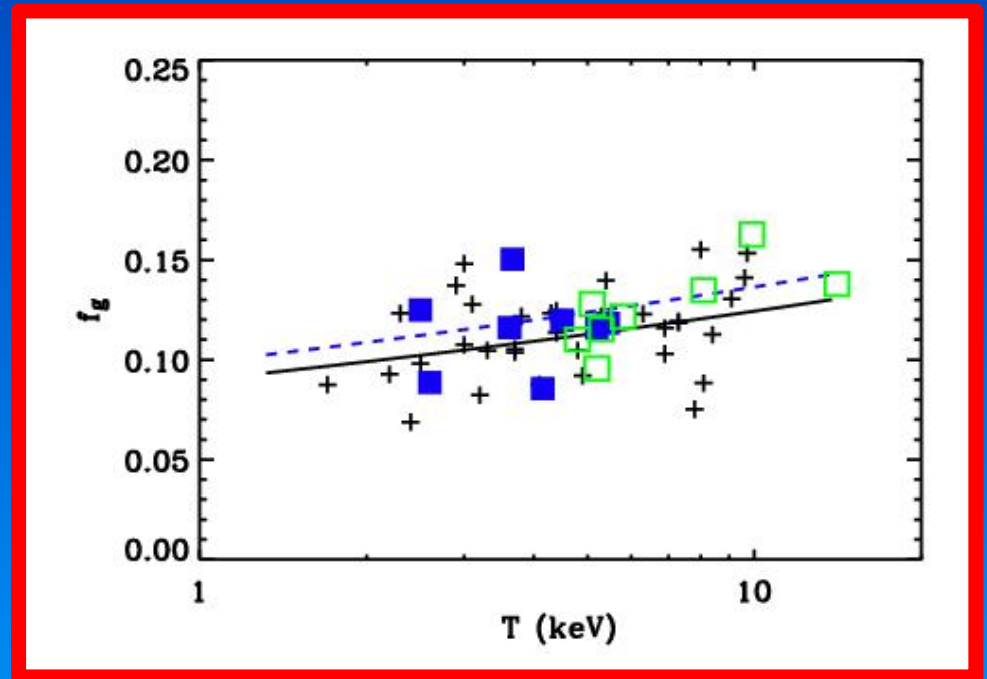
$\Delta = 1000$

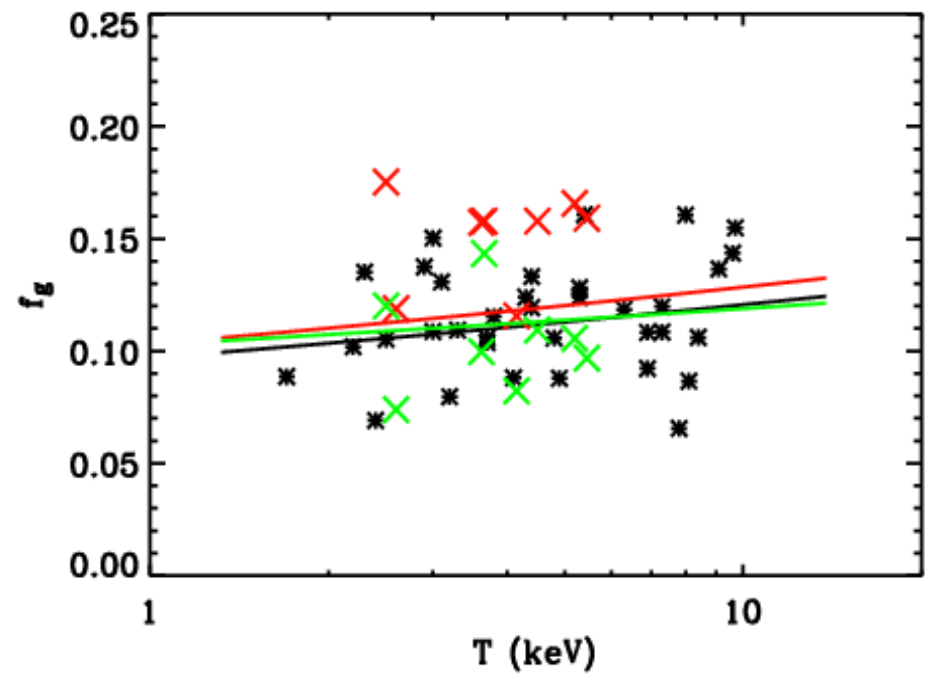
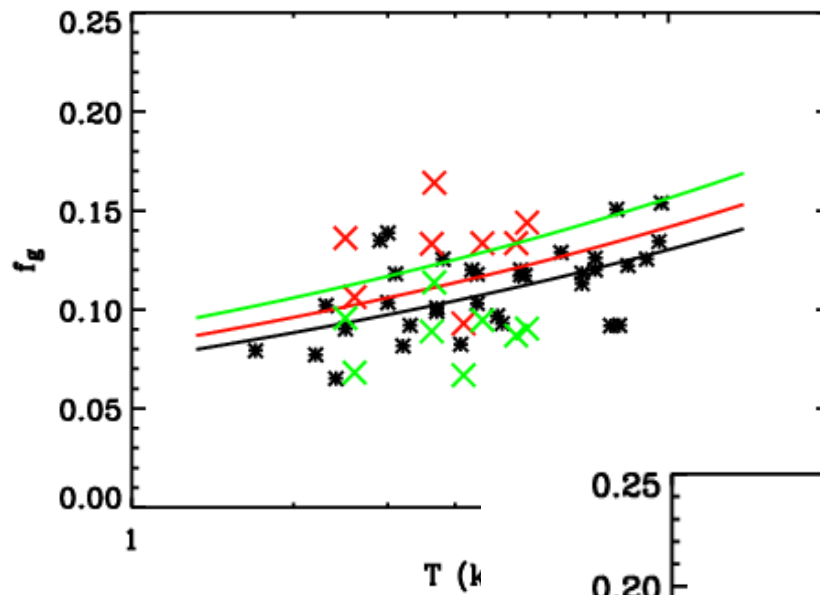
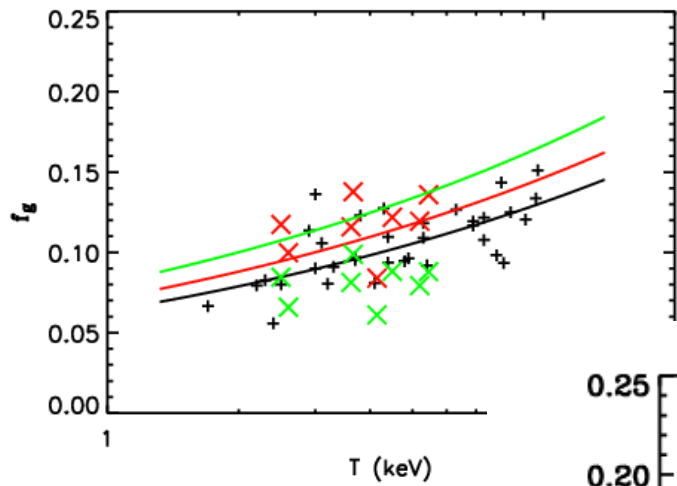


Baryon Fraction @ $z = 0.6$



$$\Delta_V$$





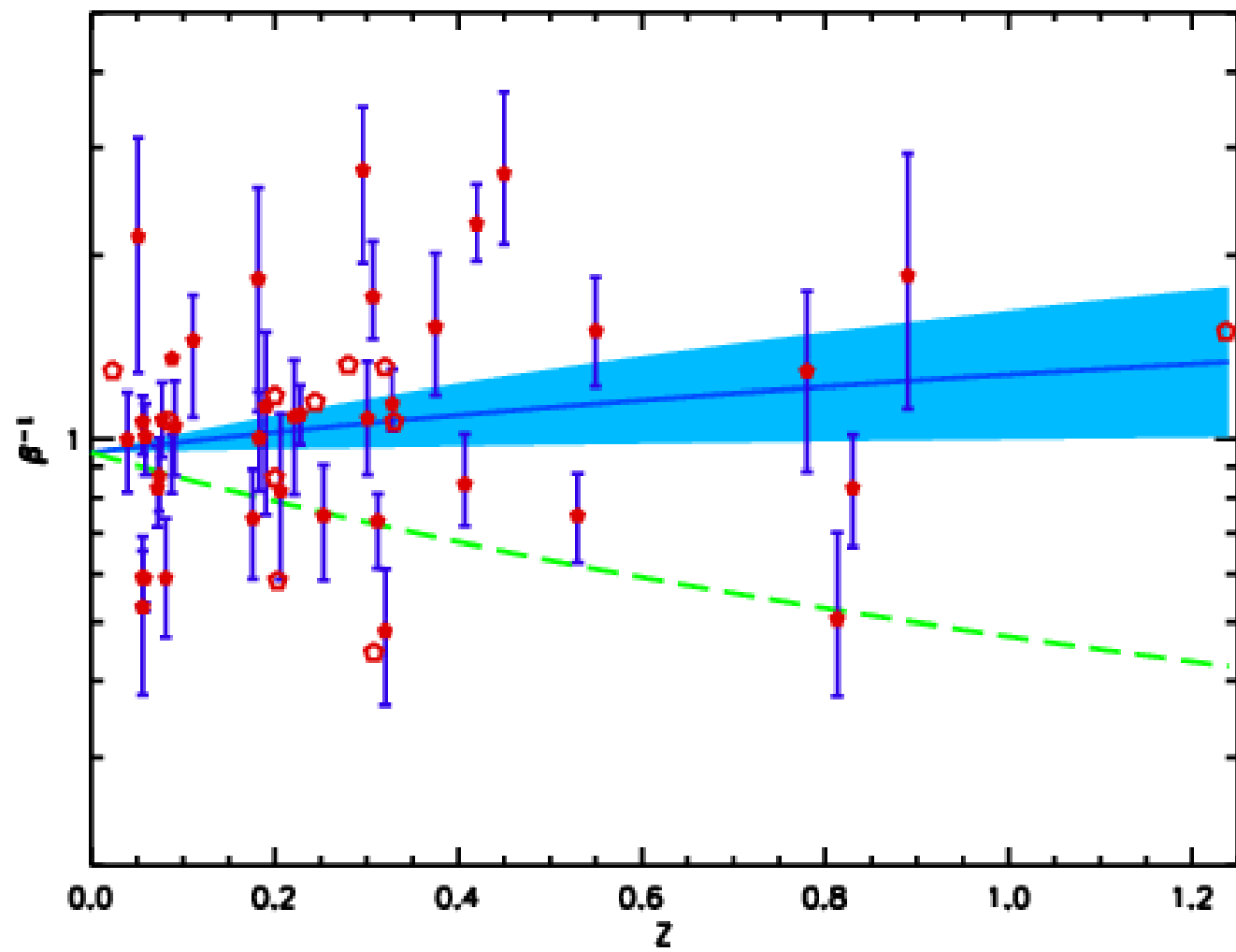
Concordance works !

Breaking the degeneracy...

$$T \propto GM/r / (1+z) \propto \sigma^2 / (1+z)$$

➤ **Testable... i.e.**

$$\beta^{-1} \propto T/\sigma^2 \propto 1/(1+z)$$



Conclusions I

- **Strong Evolution** in the abundance of x-ray clusters appears from all existing surveys in a very **consistent** way.
- **This is inconsistent with standard scaling laws in concordance model...** 😊
- **Consistent with f_b evolution ...** 😊
- **Consistent with f_b amplitude ...** 😊

Conclusions II

➤ **This could require a major revision of standard scaling of M-T (z)**
i.e. $T_x \neq GM/r$
New cluster (astro-)physics ?

➤ **No sign of it in observed clusters...**